

SCALE 1:502 000 (1 mm = 502 m) AT 250° E (110° W) LONGITUDE  
TRANSVERSE MERCATOR PROJECTION

CONTOUR INTERVAL 250 METERS  
DASHED LINES REPRESENT 125-METER CONTOUR LINES  
Planetocentric latitude and east longitude coordinate system shown in black.  
Planigraphic latitude and west longitude coordinate system shown in red.

**NOTES ON BASE**

This map, compiled photogrammetrically from Viking Orbiter stereo image pairs, is part of a series of topographic maps of areas of special scientific interest on Mars.  
MTM 500k 10/257E OMKT: Abbreviation for Mars Transverse Mercator; 1:500,000 series; center of sheet latitude 10° N, longitude 257.5° E, in planetocentric coordinate system (this corresponds to 10/102; latitude 10° N, longitude 102.5° W, in planetographic coordinate system); orthophotomosaic (OM) with color-coded (K) topographic contours and nomenclature (T) [Greeley and Batson, 1990]

**ADOPTED FIGURE**

The figure of Mars used for the computation of the map projection is an oblate spheroid (flattening of 1/176.875) with an equatorial radius of 3396.6 km and a polar radius of 3376.8 km (Kirk and others, 2002). A secondary grid (printed in red) has been added to the map as a reference to the west longitude/planetographic latitude system that is also allowed by IAU/IAAG standards (Seidelmann and others, 2002) and has been used for previous Mars maps.

**PROJECTION**

The projection is part of a Mars Transverse Mercator (MTM) system with 20° wide zones. For the area covered by this map sheet, the central meridian is at 250° E. (110° W.). The scale factor at the central meridian of the zone containing this quadrangle is 0.9960 relative to a nominal scale of 1:500,000.

**COORDINATE SYSTEM**

Longitude increases to the east and latitude is planetocentric as allowed by IAU/IAAG standards (Seidelmann and others, 2002) and in accordance with current NASA and USGS standards (Duxbury and others, 2002). A secondary grid (printed in red) has been added to the map as a reference to the west longitude/planetographic latitude system that is also allowed by IAU/IAAG standards (Seidelmann and others, 2002) and has been used for previous Mars maps.

**CONTROL**

Horizontal and vertical control was established using the Mosaic Digital Image Model 2.0 (MDIM 2.0; Kirk and others, 2000) and MOLA data. A portion of MDIM 2.0 covering the mapping area was extracted in simple cylindrical projection. This MDIM image was georeferenced to the MOLA data with an affine transformation. The MDIM image and georeferencing information were imported into a digital photogrammetric workstation (Miller and Walker, 1993) and used as an orthophoto to provide horizontal control to stereopairs of Viking imagery. The horizontal information was used to extract vertical control from the MOLA data. Note that the distribution of Viking Orbiter images suitable for mapping at a scale of 1:500,000 is uneven. Areas mapped in this series are chosen, often in blocks of two or more adjacent quadrangles, based on scientific interest as well as on the availability of suitable data for accurate mapping.

**CONTOURS**

Contours were derived from a digital terrain model (DTM) compiled on a digital photogrammetric workstation using Viking Orbiter stereo image pairs with orientation parameters derived from an analytic aerotriangulation. Contours were drawn automatically using a commercial geographic information system (GIS) software package (Environmental Systems Research Institute, 1994). For the stereomodels, the local expected vertical precision, based on image resolutions, parallax-to-height ratio (that is, convergence angle), and a matching accuracy of 0.2 pixel ranges from 43 m to 80 m, with a mean of 62 m. Elevation (in meters) is

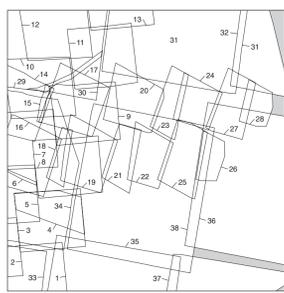
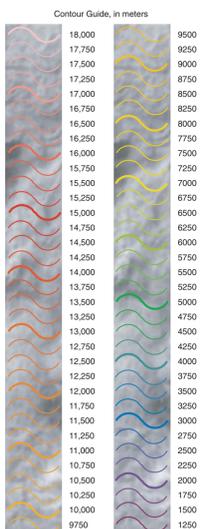
given with respect to the adopted Mars topographic datum (see "Adopted Figure" section). A comparison of the DTM values at the MOLA point locations shows that the DTM is on average 0.5 m lower than the MOLA points (n=247,122;  $\mu=0.5$  m;  $\sigma=17$  m). Contour lines were generated automatically using GIS software and were not edited. Because the contour lines were not edited, small closed contour lines, contour lines that intersect, and contour lines that do not match features are present. The post spacing for the DTM is 600 m; features that are less than 600 m in size will not be resolved and features that are smaller than 1800 m in size may only have four elevation measurements associated with them. This lack of elevation measurements may result in contour lines that do not adequately represent some features. The purpose of this mapping project is to produce the digital orthophoto and DTM. This map provides a graphical representation of the digital products that are available.

**IMAGE BASE**

The image base for this map employs Viking Orbiter images from orbits 090, 055, 892, 643, and 210. An orthophotomosaic was created on the digital photogrammetric workstation using the DTM compiled from stereo models. Integrated Software for Imagers and Spectrometers (ISIS; Torson and Becker, 1997) provided the software to project the orthophotomosaic into the Transverse Mercator Projection.

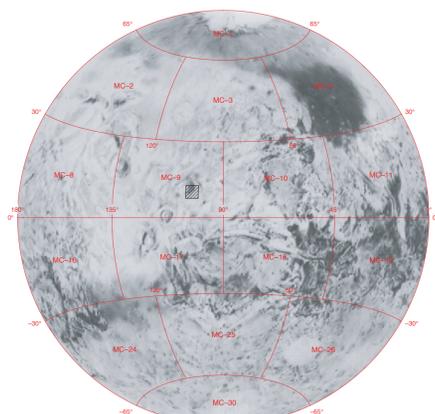
**REFERENCES**

Duxbury, T.C., Kirk, R.L., Archinal, B.A., and Neumann, G.A., 2002, Mars Geodesy/ Cartography Working Group Recommendations on Mars Cartographic Constants and Coordinate Systems, in Joint International Symposium on Geospatial Theory, Processing and Applications, Ottawa, Canada, 2002, Commission IV, Working Group 9—Extraterrestrial Mapping, Proceedings: Ottawa, Canada, International Society for Photogrammetry and Remote Sensing [http://www.isprs.org/commission4/proceedings/paper.html].  
Environmental Systems Research Institute, 1994, Arc commands: Redlands, Calif., Environmental Systems Research Institute, Inc.  
Greeley, Ronald, and Batson, R.M., 1990, Planetary mapping: New York, Cambridge University Press, p. 261–276.  
Kirk, R.L., Lee, E.M., Sucharski, R.M., Ritchie, J., Greco, A., and Castro, S.K., 2000, MDIM 2.0—A revised global digital image mosaic of Mars, in Lunar and Planetary Science XXXI: Houston, Lunar and Planetary Institute, abstract 2011 [CD-ROM].  
Miller, S.B., and Walker, A.S., 1993, Further developments of Leica Digital Photogrammetric Systems by Helava, ACSM/ASPRS Annual Convention and Exposition, Technical Papers, v. 3, p. 256–263.  
Seidelmann, P.K. (chair), Ahlkin, V.K., Bursa, M., Davies, M.E., De Bergh, C., Lieske, J.H., Oberst, J., Simon, J.L., Standish, E.M., Stooke, P., and Thomas, P.C., 2002, Report of the IAU/IAAG Working Group on Cartographic Coordinates and Rotational Elements of the Planets and Satellites: 2000: Celestial Mechanics and Dynamical Astronomy, v. 82, p. 83–110.  
Smith, D.E., Zuber, M.T., Frey, H.V., Garvin, J.B., Head, J.W., Muhlemann, D.O., Pettengill, G.H., Phillips, R.J., Solomon, S.C., Zwalby, H.J., Banerth, W.B., Davbury, T.C., Gubins, M.P., Lemoine, F.G., Neumann, G.A., Rowlands, D.D., Aharonson, O., Ford, P.G., Ivanov, A.B., McGovern, P.J., Abshire, J.B., Afzal, R.S., and Sun, X., 2001, Mars Orbiter Laser Altimeter (MOLA)—Experiment summary after the first year of global mapping of Mars, Journal of Geophysical Research, v. 106, p. 23 689–23 722.  
Torson, J.M., and Becker, K.J., 1997, ISIS—A software architecture for processing planetary images (abs.), in Lunar and Planetary Science Conference XXXVII: Houston, Lunar and Planetary Institute, p. 1443.



The following is a list of image pairs used to produce the topographic information for this map. Numbers below correspond to the numbers in the diagram above. Shaded areas indicate MOLA data.

ID	IMAGE PAIR	ID	IMAGE PAIR
1	090A43/055A23	20	210A66/210A16
2	090A44/055A23	21	210A67/210A15
3	090A45/055A23	22	210A67/210A17
4	090A46/055A23	23	210A68/210A16
5	090A46/055A23	24	210A68/210A18
6	090A46/055A25	25	210A69/210A17
7	090A47/055A25	26	210A69/210A19
8	090A48/055A25	27	210A70/210A18
9	090A48/055A25	28	210A70/210A20
10	090A50/055A25	29	892A11/643A75
11	090A51/055A27	30	892A11/643A77
12	090A52/055A27	31	892A13/643A77
13	090A53/055A27	32	892A11/643A77
14	210A12/210A62	33	892A31/643A78
15	210A62/210A14	34	892A32/643A78
16	210A62/210A13	35	892A33/643A77
17	210A64/210A14	36	892A34/643A77
18	210A65/210A13	37	892A35/643A77
19	210A65/210A15	38	892A36/643A77



Photomosaic showing location of map area. An outline of 1:5,000,000-scale quadrangles is provided for reference.

**Topographic Map of the Southeast Ascraeus Mons Region of Mars**  
MTM 500k 10/257E OMKT  
By  
U.S. Geological Survey  
2004