

NOTES ON BASE

This is one photomosaic in a set of topographic map sheets covering areas of special interest on Mars at nominal scales of 1:1,000,000 and 1:500,000 (Barton, 1973, 1975). The major source of map data was the Viking 1 spacecraft.

ADOPTED FIGURE

The figure of Mars used for the computation of the map projection is an oblate spheroid (flattening of 1/193) with an equatorial radius of 3392.4 km and a polar radius of 3375.7 km. This is not the height datum defined below under the heading "Contours."

PROJECTION

The transverse Mercator projection is used for this sheet, with a scale of 1:1,000,000. Longitude increases to the west in accordance with the usage of the International Astronomical Union (IAU, 1971). Latitudes are areographic (de Vasconcelos and others, 1973). The first meridian passes through the crater Ary-O (lat. 5, 19° S) within the crater Ary.

CONTROL

Planimetric control was derived from the primary network (Davies, 1973). A set of nine Mariner 9 pictures was tied to the primary net. The positions of Viking 1 pictures were controlled by the Mariner 9 images. No simple statement is possible for the precision, but local consistency is about 2 km.

IMAGE PROCESSING

Viking 1 frames were specially processed in the computer and mosaicked. Processing included artifact and noise removal, contrast enhancement, and spatial filtering to remove camera shading and to enhance fine detail in the image. Processing of all frames except those in the northwest corner was done by the Mission Test and Video System (MTVS) at the Jet Propulsion Laboratory (JPL).

CONTOURS

Since Mars has no sea and hence no sea level, the datum (the 0 km contour line) for altitudes is defined by a gravity field described by spherical harmonics of fourth order and fourth degree (Jordan and Lowell, 1973) combined with a 9 millibar atmospheric pressure surface derived from radio-occultation data (Elve and others, 1973; Christensen, 1975; Wu, 1975).

The contour line were compiled by stereophotogrammetric methods from pairs of Viking 1 pictures taken on revolution 27. Special stereoplotter methods were applied to this extremely narrow-angle (1°) photography (Wu, 1975). Parameters for setting stereomodels in analytical plotters were derived from analytical triangulation performed with the U.S. Geological Survey GLANT computer program, the prototype of MUSAAT, (Elasol, and others, 1970). This blends photogrammetric coordinate measures with position and orientation values taken from the Supplementary Experiment Data Record (SEDR) with appropriate weights for their disparate data.

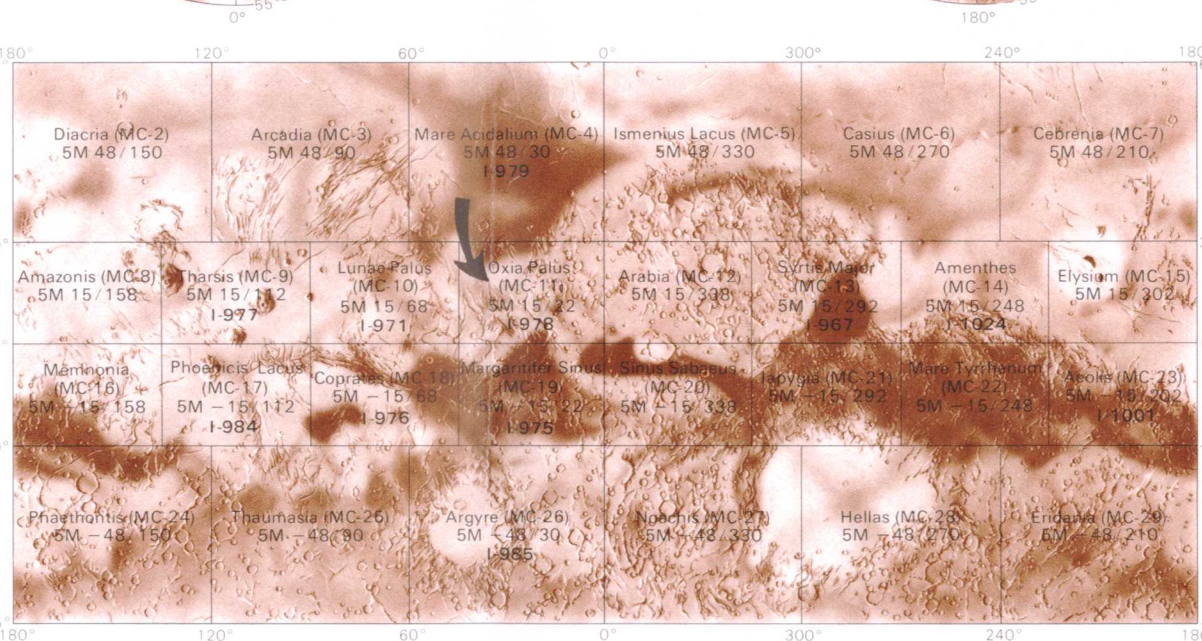
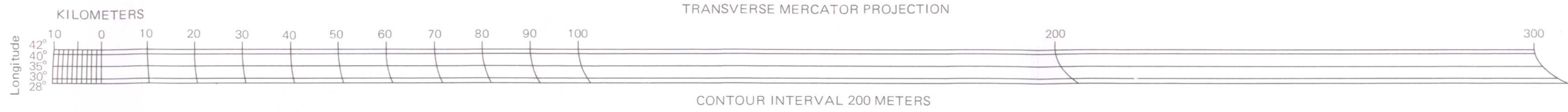
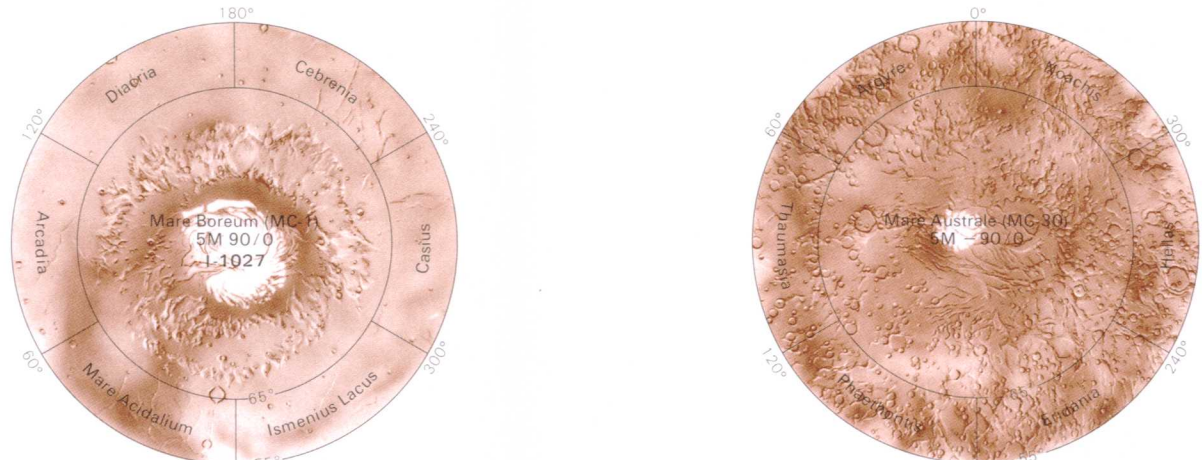
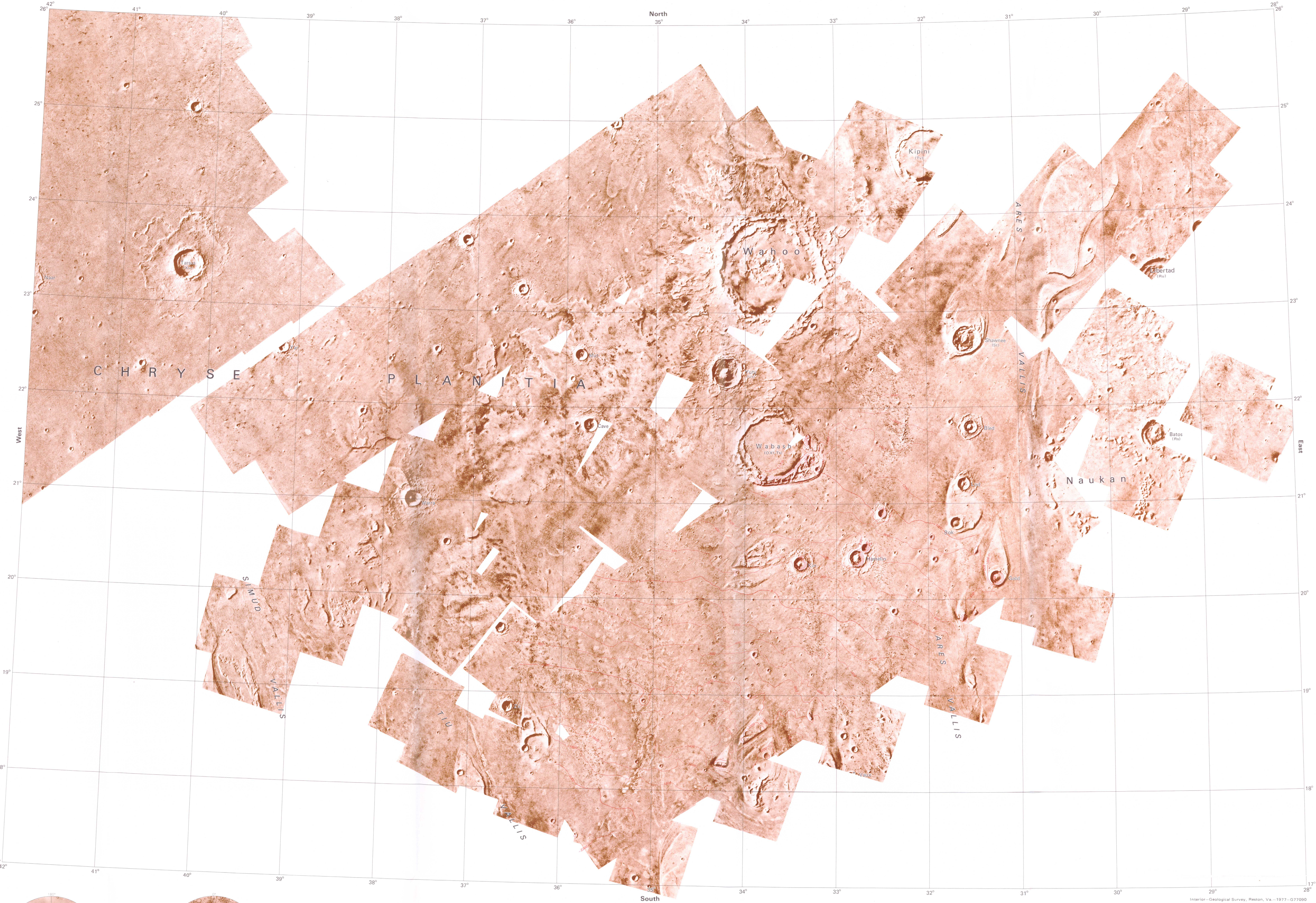
NOMENCLATURE

Names on this sheet are approved by the International Astronomical Union (1974, 1977). Named craters bearing double letters in parentheses are designated by the same letters on the 1:5,000,000 Oxis Palus sheet that covers this area. These double letter designations refer to position on the 1:5,000,000 sheet and are derived from a grid based on equidistant meridians and parallels; the alphabet (I and O omitted) runs in the direction of increasing longitude (W) and latitude (N). The complete designation of a crater in the name of the quadrangle followed by a double letter. The prefix OXI (identifying the Oxis Palus sheet) is part of the complete designation but, for brevity, is not shown on most craters.

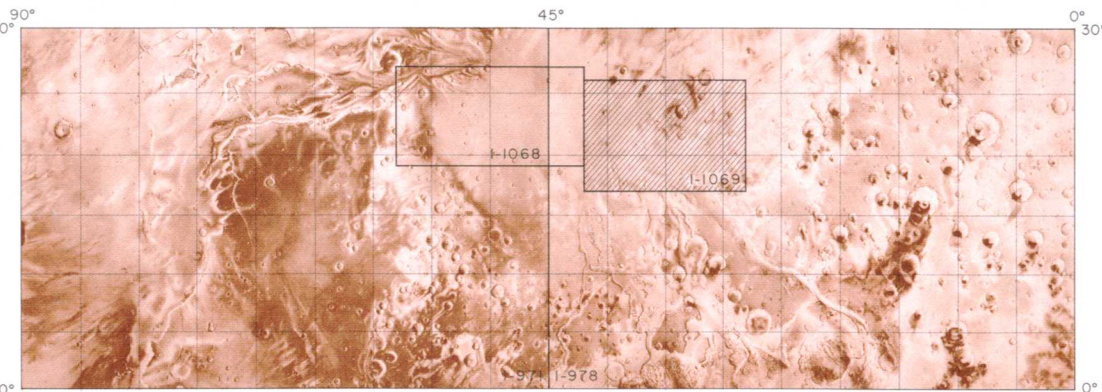
M IM 22/35 CMC: Abbreviation for Mars, 1:1,000,000 series; center of sheet, 22° N, lat 35° long; controlled mosaic, CM, with contours, C.

REFERENCES

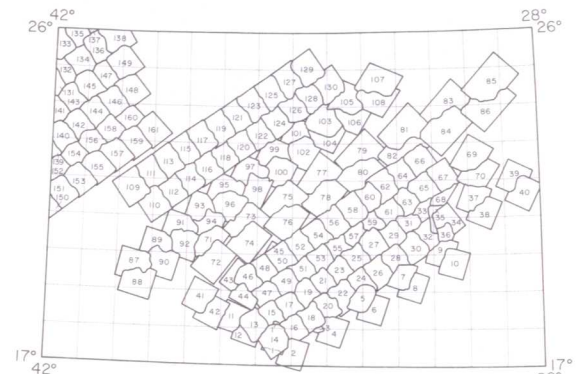
- Barton, R. M., 1973, Cartographic products from the Mariner 9 mission: Jour. Geophys. Research, v. 78, no. 20, p. 4524-4535.
—, 1976, Cartography of Mars: 1975: The American Cartographer, v. 3, no. 1, p. 57-63.
Christensen, E. J., 1975, Martian topography derived from occultation, radar, spectral, and optical measurements: Jour. Geophys. Research, v. 80, no. 20, p. 2909-2913.
Davies, M. E., 1973, Mariner 9: Primary control net: Photogram. Eng., v. 39, no. 12, p. 1297-1302.
Davies, M. E., and Arthur, D. W. G., 1973, Martian surface coordinates: Jour. Geophys. Research, v. 78, no. 20, p. 4354-4394.
Elasol, A., Brewer, R. K., Gracie, G., and Crombie, M., 1970, MUSAAT IV, Final technical report, U.S. Army Engineers Topographic Laboratories, Nov. 1970.
International Astronomical Union, Commission 16, 1971, Physical study of planets and satellites, in Proc. 14th General Assembly, 1970: Internat. Astron. Union Trans., v. XVII, p. 128-137.
—, 1974, Physical study of planets and satellites, in Proc. 15th General Assembly, 1973: Internat. Astron. Union Trans., v. XVII, p. 104-108.
—, 1977, Physical study of planets and satellites, in Proc. 16th General Assembly, 1976: Internat. Astron. Union Trans. (in press).
Jordan, J. P., and Lowell, Jack, 1973, Mariner 9 an instrument of dynamical science: Presented at AAS/AIAA Astrodynamics Conf., Vol. Colo., July 16-18, 1973.
Kliore, A. J., Priddey, Gunnar, Sedel, R. L., Sykes, M. J., and Woicakyn, P. M., 1973, S-band radio occultation measurements of the atmosphere and topography of Mars with Mariner 9: Extended mission coverage of polar and intermediate latitudes: Jour. Geophys. Research, v. 78, no. 20, p. 4313-4381.
de Vasconcelos, G. D., Davies, M. E., and Strum, F. M., Jr., 1973, The Mariner 9 areographic coordinate system: Jour. Geophys. Research, v. 78, no. 20, p. 4393-4405.
Wu, S. S. C., 1975, Topographic mapping of Mars: U.S. Geol. Survey Interagency Report: Astrogeology 63, 191 p.



NOTE TO USERS
Users noting errors or omissions are urged to indicate them on the map and to forward the map to Astrogeologic Studies, Geologic Division, 2205 North Gemini Drive, Flagstaff, Arizona 86001. A replacement copy will be returned.



MAP OF LUNAE PALUS AND OXIA PALUS QUADRANGLES
SHOWING LOCATION OF THIS MAP



INDEX TO VIKING 1 PICTURES USED ON MOSAIC

Index No.	Picture No.	Index No.	Picture No.	Index No.	Picture No.	Index No.	Picture No.	Index No.	Picture No.	Index No.	Picture No.	Index No.	Picture No.	Index No.	Picture No.
1	3449	24	4434	47	6438	70	3423	93	3417	116	3443	139	10460		
2	3450	25	4445	48	6439	71	3421	94	3418	117	3441	140	10462		
3	3451	26	4446	49	6436	72	3422	95	3419	118	3442	141	10463		
4	3452	27	4447	50	6437	73	3420	96	3420	119	3443	142	10464		
5	3453	28	4448	51	6438	74	3421	97	3421	120	3444	143	10465		
6	3454	29	4449	52	6439	75	3422	98	3422	121	3445	144	10466		
7	3455	30	4450	53	6440	76	3423	99	3423	122	3446	145	10467		
8	3456	31	4451	54	6441	77	3424	100	3424	123	3447	146	10468		
9	3457	32	4452	55	6442	78	3425	101	3425	124	3448	147	10469		
10	3458	33	4453	56	6443	79	3426	102	3426	125	3449	148	10470		
11	3459	34	4454	57	6444	80	3427	103	3427	126	3450	149	10471		
12	3460	35	4455	58	6445	81	3428	104	3428	127	3451	150	10472		
13	3461	36	4456	59	6446	82	3429	105	3429	128	3452	151	10473		
14	3462	37	4457	60	6447	83	3430	106	3430	129	3453	152	10474		
15	3463	38	4458	61	6448	84	3431	107	3431	130	3454	153	10475		
16	3464	39	4459	62	6449	85	3432	108	3432	131	3455	154	10476		
17	3465	40	4460	63	6450	86	3433	109	3433	132	3456	155	10477		
18	3466	41	4461	64	6451	87	3434	110	3434	133	3457	156	10478		
19	3467	42	4462	65	6452	88	3435	111	3435	134	3458	157	10479		
20	3468	43	4463	66	6453	89	3436	112	3436	135	3459	158	10480		
21	3469	44	4464	67	6454	90	3437	113	3437	136	3460	159	10481		
22	3470	45	4465	68	6455	91	3438	114	3438	137	3461	160	10482		
23	3471	46	4466	69	6456	92	3439	115	3439	138	3462	161	10483		

CONTROLLED MOSAIC OF THE EASTERN PART OF THE CHRYSE PLANITIA REGION OF MARS

M IM 22/35 CMC

1977