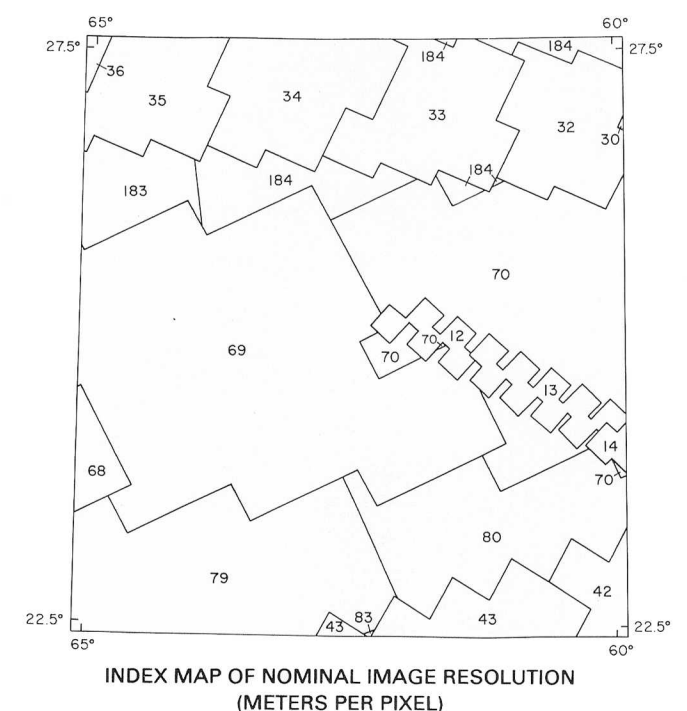


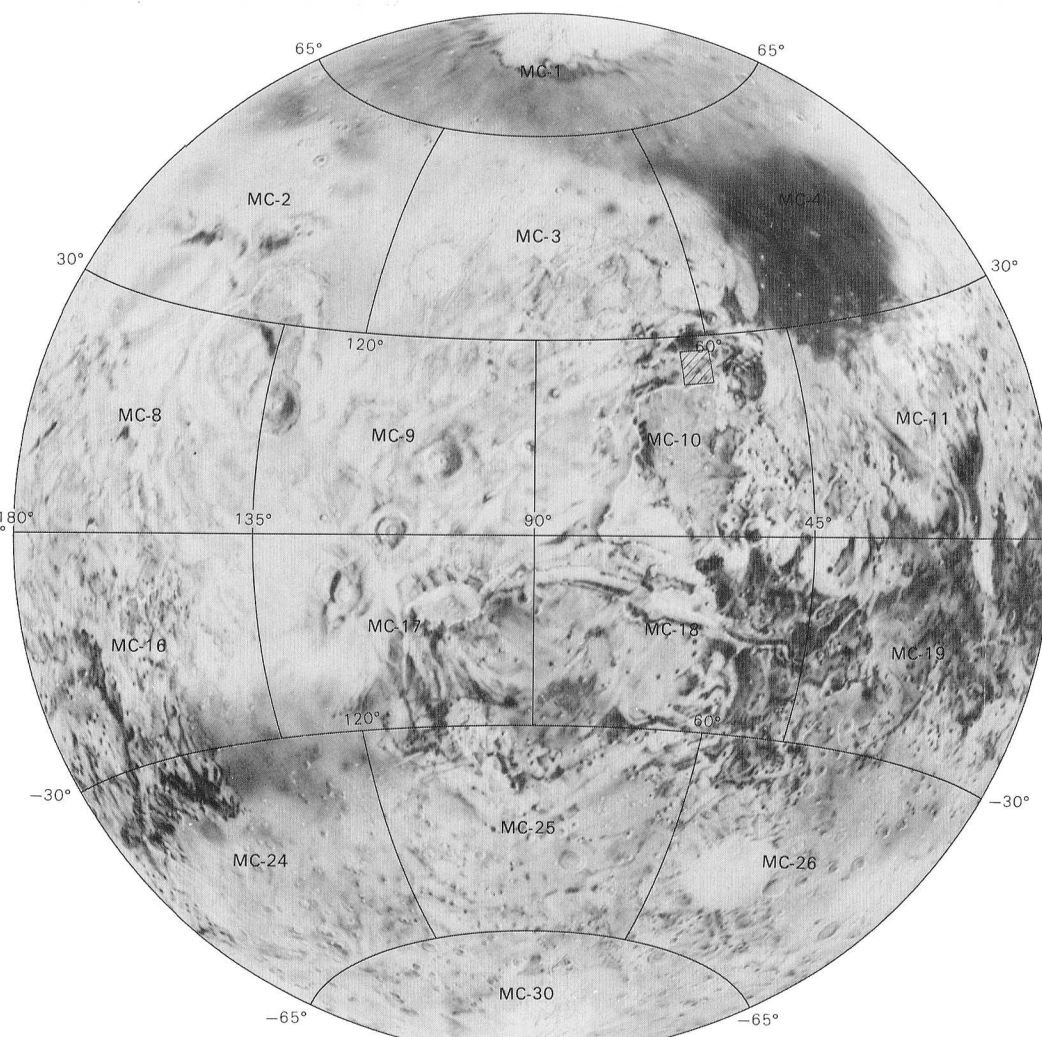
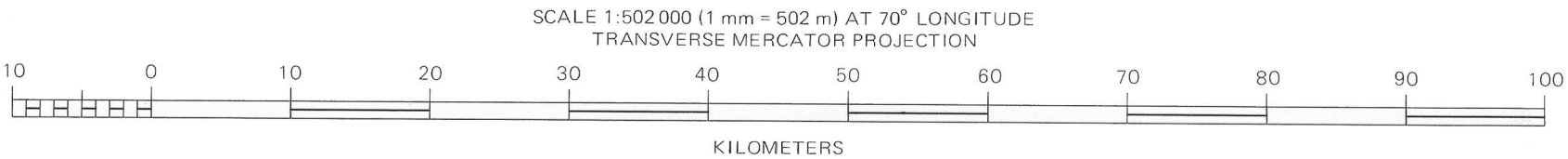
The mosaic was made with the Viking images outlined above. Copies of various enhancements of these images are available from National Space Science Data Center, Code 601, Goddard Space Flight Center, Greenbelt, MD 20771.



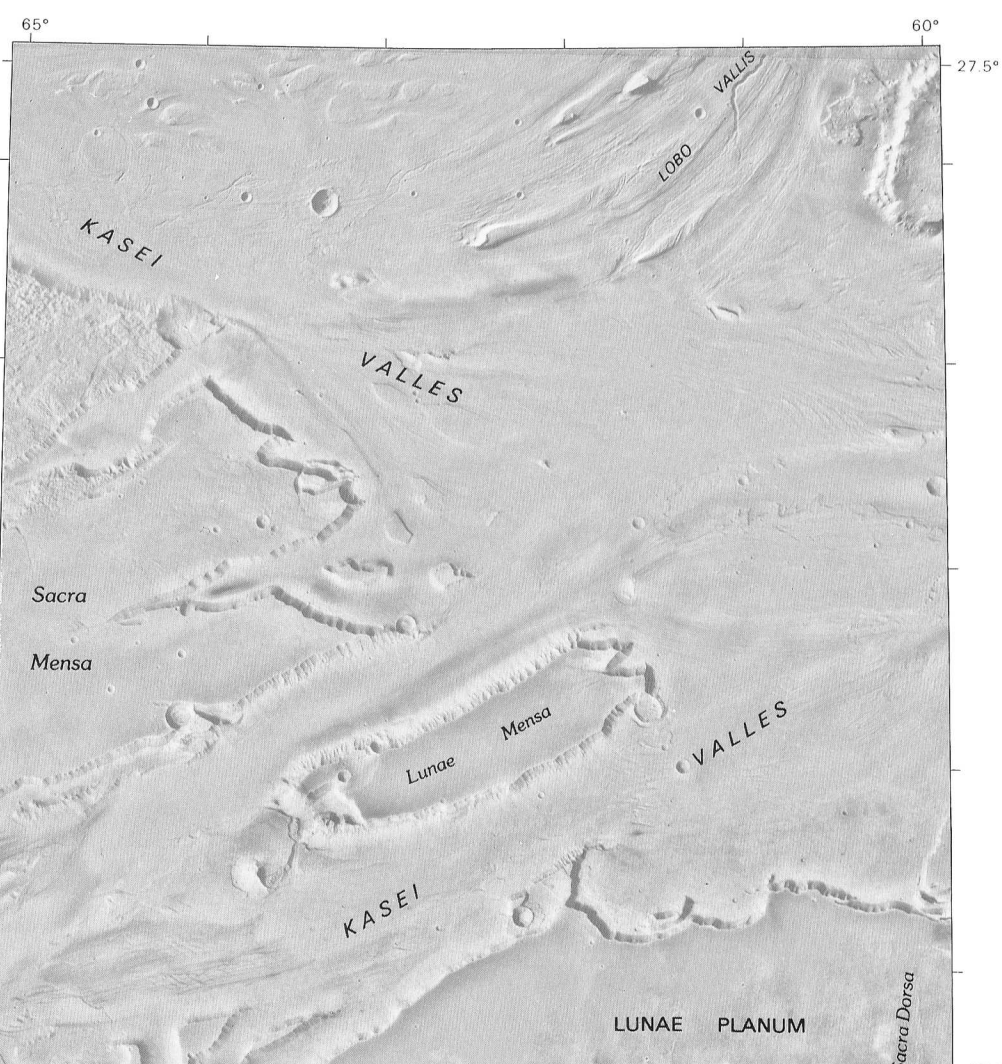
NOTES ON BASE
This photomosaic covers part of an area of special scientific interest on Mars. It is published in a series designed to support topical studies, which is not expected to result in systematic coverage of the planet. The mosaic was compiled by digital methods described by Batson (1987) and Edwards (1987).
The distribution of Viking Orbiter images suitable for mapping at a scale of 1:500,000 is uneven, as are the quality and distribution of map controls. The mosaics are usually compiled in blocks of two or more adjacent quadrangles that are selected on the basis of scientific importance, not necessarily in areas of optimum coverage by high-resolution images or precise map controls. Image placement is based on a planetwide topographic control net that has a published standard error of 5 km (Wu and Schaler, 1984). A block of mosaics compiled in an area where controls have optimum distribution and precision is not likely to match adjacent blocks previously compiled in areas where controls are sparse or imprecise. Where discrepancies exist between adjacent mosaics, the more recent compilation is likely to be more accurate.
The projection is part of a Mars Transverse Mercator (MTM) system with 20° zones. The scale factor at the central meridian of the zone containing this quadrangle is 0.9960. The projection scale is based on an oblate spheroid (flattening of 1/192) with an equatorial radius of 3393.4 km and a polar radius of 3370.7 km.
Digital processing and mosaicking were done by Patricia K. Thomas.

NOMENCLATURE
All names shown on the reduced base mosaic are approved by the International Astronomical Union (IAU), 1974, 1986, and 1989.
M 500k 25/62 CM: Abbreviation for Mars; 1:500,000 series; center of sheet lat 25° N., long 62°; controlled photomosaic (CM).

REFERENCES
Batson, R.M., 1987. Digital cartography of the planets: New methods, its status, and its future. *Photogrammetric Engineering and Remote Sensing*, v. 53, no. 9, p. 1211-1218.
Edwards, Kathleen, 1987. Geometric processing of digital images of the planets: *Photogrammetric Engineering and Remote Sensing*, v. 53, no. 9, p. 1219-1222.
International Astronomical Union, 1974. Commission 16. Physical study of planets and satellites and Lunar and martian nomenclature, in *Proceedings of the 15th General Assembly, Sydney, 1973: Transactions of the International Astronomical Union*, v. 15B, p. 105-108, 217-221.
1986. Working Group for Planetary System Nomenclature, in *Proceedings of the 19th General Assembly, New Delhi, 1985: Transactions of the International Astronomical Union*, v. 19B, p. 347-350.
1989. Working Group for Planetary System Nomenclature, in *Proceedings of the 20th General Assembly, Baltimore, 1988: Transactions of the International Astronomical Union*, v. 20B, p. 704-705.
Wu, S.S.C., and Schaler, F.J., 1984. Mars control network: American Society of Photogrammetry, in *Technical papers of the 50th annual meeting of the American Society of Photogrammetry*, v. 2, Washington, D.C., March 11-16, 1984, p. 456-463.



Photomosaic location is shown in the western hemisphere of Mars. An outline of 1:500,000-scale quadrangles is provided for reference.



Contrast in the reduced base mosaic was purposely suppressed to emphasize the names

CONTROLLED PHOTOMOSAIC OF THE MTM 25062 QUADRANGLE (REVISED), KASEI VALLIS REGION OF MARS