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A Description of Colored Gravity and Terrain
Maps of the Southwestern Cordillera

by

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INTRODUCTION

Colored gravity and terrain maps of the Southwestern Cordillera covering California, Nevada, and portions of Utah and Arizona, have been prepared using an Applicon Color plotter. Colored copies of these maps are included in this report. Standard 2" x 2" color slides of these maps are available from the U.S. Geological Survey Photo Library (Mail Stop 914, Box 25046, Denver Federal Center, Denver, CO, 80225, telephone 303/234-4004). When ordering slides from the library, refer to the open-file number of this report.

The gravity maps were prepared as part of regional studies of the Basin and Range Province and to extend the California Isostatic map (Roberts and others, 1981). The 30 maps fall into six classifications:

- 1) Digital terrain (1 map)
- 2) Wavelength filtered terrain (6 maps)
- 3) Complete Bouguer anomaly gravity (3 maps)
- 4) Isostatic residual gravity (the result of removing an isostatic regional gravity surface from the complete Bouguer anomaly, 7 maps)
- 5) Isostatic regional gravity (a smooth surface based on the calculated gravity effect of the crustal root required to provide complete local isostatic compensation of topography, 7 maps)
- 6) Wavelength filtered Bouguer gravity (6 maps)

This report supercedes Open-File Report 82-839 (Saltus, 1982) which contained only the maps shown in figures 8 and 11.

THE DATA SETS

1. Gravity Data

The gravity maps are based on about 100,000 gravity stations. This includes about 64,000 observations in California (Snyder and others, 1981a, 1981b) which had to be converted to the International Gravity Standardization Net of 1971 (Morelli, 1974) and reduced using the 1967 Geodetic Reference System formula for theoretical gravity (International Association of Geodesy, 1971) in order to agree with the Nevada data. In Nevada, data from the following ten 1° x 2° degree quadrangles were used:

- 1) Caliente (Healey and others, 1981)
- 2) Death Valley (Healey and others, 1980)
- 3) Goldfield (Healey and others, 1980)
- 4) Kingman (Bracken and Kane, 1983)
- 5) Las Vegas (Kane and others, 1979)
- 6) Lund (Bol and others, 1983)
- 7) Millett (Erwin and Bittleston, 1977)
- 8) Reno (Erwin and Berg, 1977)

- 9) Tonopah (Healey and others, 1981)
- 10) Walker Lake (D. Plouff, personal commun.).

Data for the remainder of Nevada, Utah, and Arizona are from T. G. Hildenbrand, U.S. Geological Survey, Denver, and consist mostly of Department of Defense (DOD) data available through the National Oceanic and Atmospheric Administration (NOAA) Data Center. (NOAA, National Geophysical and Solar Terrestrial Data Center, Boulder, CO 80303).

2. Terrain Data

The digital terrain data are also available from NOAA in Boulder. The data are a subset of 5-minute x 5-minute average elevation data covering North America. The data are described in U.S. Department of Commerce Data Announcement 1980 SE-U.

A FORTRAN gridding program (Webring, 1981), based on minimum curvature was used to interpolate and extrapolate both the gravity and terrain data to a square grid with a 5 x 5 km grid interval. The gravity and terrain data were projected using the Albers Conic Equal-Area projection with standard parallels for the U.S. (latitude 29.5°N, latitude 45.5°N). A central meridian of longitude 96°W was used. The gridded data were plotted using Applicon Incorporated proprietary color software.

DESCRIPTION OF THE MAPS

Digital Terrain Map of the Southwestern Cordillera (Figure 1)

The NOAA 5 minute data set was contoured at 250 meter intervals. The color contours range from -250 to 4,000 meters. Blue colors represent the lower elevations and red colors represent the higher elevations.

Wavelength Filtered Terrain Maps of the Southwestern Cordillera (6 maps, Figures 2-7)

A fast Fourier transform filtering program by Robert Simpson, USGS, Menlo Park (written communication) was used to select long wavelengths from the digital terrain data. The long wavelength data were subtracted from the terrain map to produce the short wavelength map. This low pass filter used a simple rectangular window in the frequency domain, modified so that the gain of the filter drops from one to zero along a ramp centered at the cut-off wavelength. Three sets of filtered maps were made with wavelength cut-offs of 100, 150, and 250 km. The parameters for each map are listed in table 1.

Table 1

Figure	2	3	4	5	6	7
wavelengths km	>100	>150	>250	<100	<150	<250
linear ramp km	80-120	125-175	200-300	80-120	125-175	200-300
contour interval meters	250	250	250	125	125	125

For a discussion of wavelength filtering see Simpson and others (1982).

Bouguer Gravity Anomaly Maps of the Southwestern Cordillera (3 maps, Figures 8, 9, and 10)

Bouguer anomalies were calculated using the Geodetic Reference System 1967 formula (International Association of Geodesy, 1971) and tied to the IGSN 71 datum (Morelli, 1974) with reduction densities of 2.67, 2.5, and 2.3 g/cm³. Terrain corrections have been calculated using a program by Plouff (1977). Estimated error in the Bouguer gravity value is less than 2.0 milligals (mGal). The terrain corrections were calculated with a Bouguer reduction density of 2.67 g/cm³ and then converted to 2.5 g/cm³ and 2.3 g/cm³ using the formula

$$TC_{p2} = TC_{2.67} \times \frac{p2}{2.67} ,$$

(where p2 is the alternate Bouguer density, TC_{p2} is the terrain correction for Bouguer density p2 and TC_{2.67} is the terrain correction for density 2.67 g/cm³). This method is not correct for stations near the ocean, as terrain corrections for the variation of the ocean bottom from sea level were made using a density of 1.67 g/cm³ (2.67 g/cm³ minus the density of sea water, 1.0 g/cm³). For these areas the correct conversion would be

$$TC_{p2} = TC_{ocean} \times \frac{(p2 - 1.0)}{1.67} .$$

The error for a large oceanic terrain correction, e.g. 20 mGal, would be 1.66 mGal. Since the correction is added to the gravity value, this means that coastal Bouguer values will be slightly high. This inaccuracy is too small, however, to substantially alter the 20 mGal contours, even at the shore line.

The color contours of the maps range from -280 mGal to 100 mGal at an interval of 20 mgal, with blue colors representing areas of low Bouguer gravity values and red colors showing areas of high Bouguer gravity values.

Isostatic Residual Gravity Maps of the Southwestern Cordillera (7 maps, Figures 11-17)

Each of these maps is the result of subtracting an isostatic regional surface (figures 18-24) from a complete Bouguer anomaly map (figures 8-10).

Isostatic corrections were computed using a FORTRAN program (Simpson, written communication) based on the Airy-Heiskanen model of compensation (Heiskanen and Meinesz, 1958, p. 135-137) using the parameters shown in table 2. The isostatic residual values range from -70 mGal to +70 mGal with a contour interval of 10 mgal. Blue colors represent areas of low residual gravity and reds show areas of high residual gravity. The reference to Moho in the figure captions refers simply to the bottom of the root. This surface will not necessarily agree with the seismically determined Moho.

Table 2

Residual Map figure	11	12	13	14	15	16	17
Regional Map figure	18	19	20	21	22	23	24
Crustal Density: g/cm ³	2.67	2.67	2.67	2.67	2.67	2.5	2.3
Compensation Depth: km	25	25	25	25	30	25	25
Density Contrast at compensation depth: g/cm ³	0.4	0.2	0.3	0.6	0.4	0.4	0.4

Isostatic Regional Maps of the Southwestern Cordillera (7 maps, Figures 18-24)

These maps show the gravitational attraction of the crustal root required by the model to provide complete local isostatic compensation of topography.

Isostatic corrections were computed using the parameters shown in table 2. The isostatic regional values are contoured with the same color scale as the Bouguer maps (range -280 to +100 mGal, 20 mGal interval).

Wavelength Filtered Bouguer Gravity Maps of the Southwestern Cordillera (6 maps, Figures 25-30)

Three sets of wavelength filtered Bouguer maps were made using the same procedure as for the filtered terrain maps. The longer wavelength maps are contoured using the same color scale as the Bouguer maps (range -280 to +100 mGal, 20 mGal interval). The short wavelength maps are contoured using the same color scale as the isostatic residual maps (range -70 to +70 mGal, 10 mGal interval). The parameters for each map are shown in Table 3. For a discussion on effects of wavelength filtering see Simpson and others (1982).

Table 3

Figure	25	26	27	28	29	30
wavelength cutoff km	>100	>150	>250	<100	<150	<250
linear ramp km	80-120	125-175	200-300	80-120	125-175	200-300
contour interval mGal	20	20	20	10	10	10

Discussion and Interpretation

This series of maps was inspired by a desire to compare the results of wavelength filtering vs. isostatic modeling in the analysis of the Bouguer gravity map. The wavelength filter is a somewhat arbitrary mathematical surface fitting process, while the isostatic regional results from a model of compensation designed to explain physical observations. That the two methods are comparable is evident from the basic similarity of figures 18 (isostatic regional density = 2.67 g/cm³, contrast = 0.4 g/cm³, compensation depth = 25 km) and 27 (filtered Bouguer gravity containing only wavelengths longer than 250 km). This correlation indicates the extent to which the gravitational effect of the Airy root accounts for wavelengths longer than 250 km of the Bouguer gravity field.

However, there are important and interesting differences between the two methods. Although the isostatic regional (figure 18) resembles the >250 km wavelength map (figure 27), it has areas which contain shorter wavelength

features as well. For example, the shape of the regional field in north-central California looks more like the same area of the >150 km wavelength map (figure 26). Another example is southernmost California where the high coming in from Mexico on the regional map (figure 18) better resembles the same area on the >100 km wavelength map (figure 25). The central Nevada low is not as deep in the regional map as in any of the long wavelength maps.

The differences are even more striking on the residual maps. Compare the <250 km wavelength map (figure 30) with the isostatic residual map (figure 11). One of the most striking features of the <250 km wavelength map is the series of parallel alternating highs and lows that extend east of north from the California coast to the Nevada-Utah state line. Except for the westernmost low and high, these stripes have been interpreted (Kane and others, 1982) as parallel bands of felsic rock causing gravity lows and intervening bands of normal crust causing relative highs. Kane comments further (Kane and Godson, 1983) that the strike of these features parallel to the subducted oceanic plate prompts speculation that the pattern is related in some way to subduction.

These stripes are not so obvious on the isostatic residual map (figure 11). The westernmost low and high (the great valley low and the high associated with the west side of the Sierra Nevada) are still clear, although the north-south continuity of the high is rather vague. The remaining stripes which cross Nevada appear only as a residual of the symmetric central Nevada "Butterfly" anomaly (Eaton and others, 1978, p. 78, Fig. 3-10) and do not show as much north-south extent as the stripes on the wavelength filtered map. Thus, based on the isostatic residual map one would be unlikely to postulate some relation between oceanic subduction off the California coast and the faint linearities suggested by the center and sides of the "Butterfly".

The isostatic map (figure 11) shows a prominent low in northwestern California which extends from Cape Mendocino east to the great valley. This low has been interpreted to "reflect the south edge of the Gorda plate where it is subducted eastward beneath the North American plate" (Jachens and Griscom, 1983). The proximity at the coast of the western edge of this gravity low to magnetic and seismic expressions of the plate in that region led Jachens and Griscom to associate this anomaly with the Gorda plate. On the <250 km wavelength map, however, the low does not extend all the way to the coast. In addition, the eastern-most lobe of the low as shown on isostatic map (fig. 11) is missing on the <250 km wavelength map (figure 30). The use of the wavelength residual technique over the isostatic method could have led to a different interpretation in this case, just as in the case of the parallel stripe features previously mentioned.

In short, although many of the same basic features are enhanced by these two residual techniques, the differences in the residual maps can lead to very different interpretations. Because the isostatic technique is based on a physical model it is perhaps safer to attribute geologic interpretations to its anomalies. Thus, the isostatic residual map is probably a more useful starting point for quantitative modelling. However, care must be used in attaching significance to either model. Despite the stability of this isostatic model, it is still only a model with a uniform density contrast, crustal density, and compensation depth. Clearly, true isostatic compensation does not follow such simple rules. The wavelength filtering technique makes the assumption that the gravitational effects of near-surface masses (generally

those of geologic interest) can be separated from those arising from larger and deeper features (such as the isostatic root system, whatever its structure may be) simply by separating wavelengths. This assumption, too, is clearly naive. The separation is not complete, some long wavelength anomalies can be caused by broad shallow features. Also, the amplitudes of short wavelength anomalies may be distorted by the removal of the long wavelengths.

Examination of the isostatic regional maps and their associated residual maps reveals the stability of the Airy model: despite variation of the model parameters to either end of a spectrum of reasonable crustal densities and density contrasts, the overall appearance of the residual map is preserved. The maximum change in value is about 20 mGal in the depth of the central Nevada low. Therefore, uncertainties in the exact parameters for the isostatic model do not preclude its general usefulness.

The inverse correlation of wavelength filtered terrain (especially figure 4, with wavelengths longer than 250 km), and the isostatic regional (figure 18) is inherent in the basic concept of isostatic compensation (the topographic surface load is in some sort of hydrostatic equilibrium). In this case, the shape of the supporting Airy root is based on the surface topography, and therefore, the isostatic regional shows some of the same long-wavelength features as the wavelength filtered terrain maps (figures 2-4).

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Figure 1

DIGITAL TERRAIN MAP OF THE SOUTHWESTERN CORDILLERA

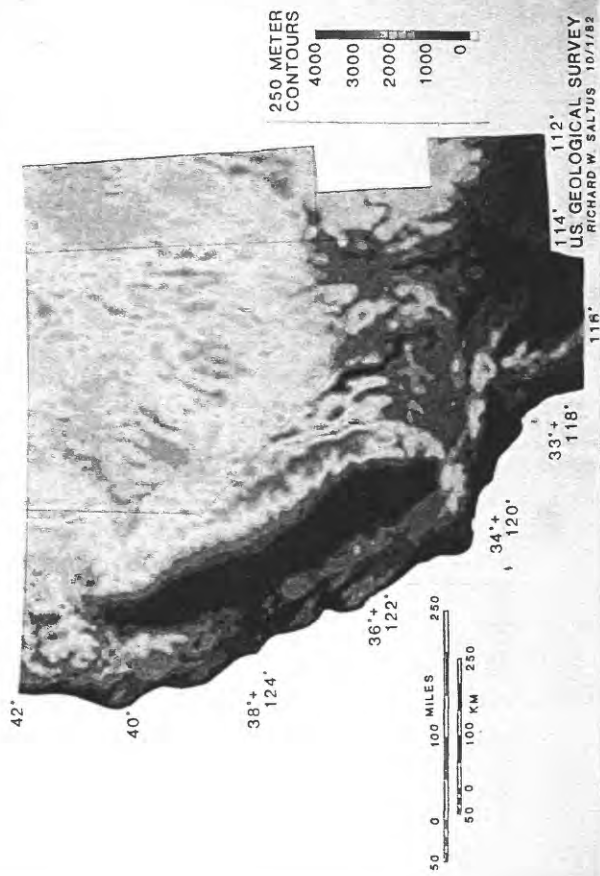
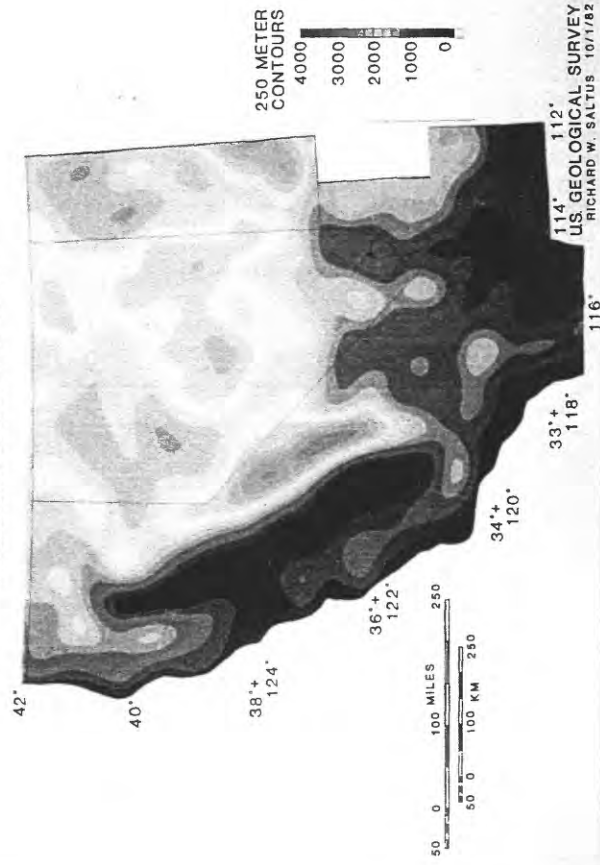


Figure 2

USGS Open-file 84-95

WAVELENGTH FILTERED TERRAIN MAP OF THE SOUTHWESTERN CORDILLERA
WAVELENGTHS LONGER THAN 100 KM



WAVELENGTH FILTERED TERRAIN MAP OF THE SOUTHWESTERN CORDILLERA
WAVELENGTHS LONGER THAN 150 KM

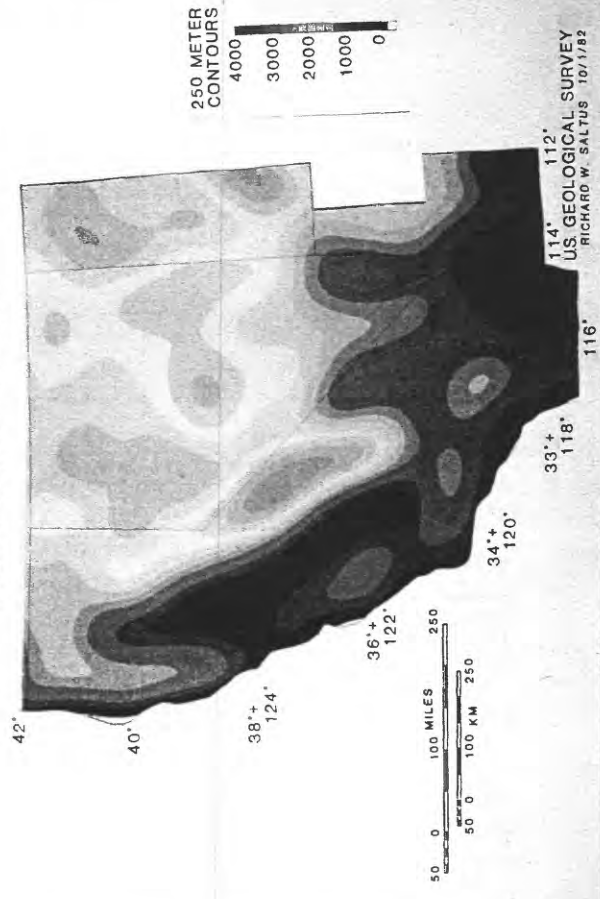


Figure 3

WAVELENGTH FILTERED TERRAIN MAP OF THE SOUTHWESTERN CORDILLERA
WAVELENGTHS LONGER THAN 250 KM

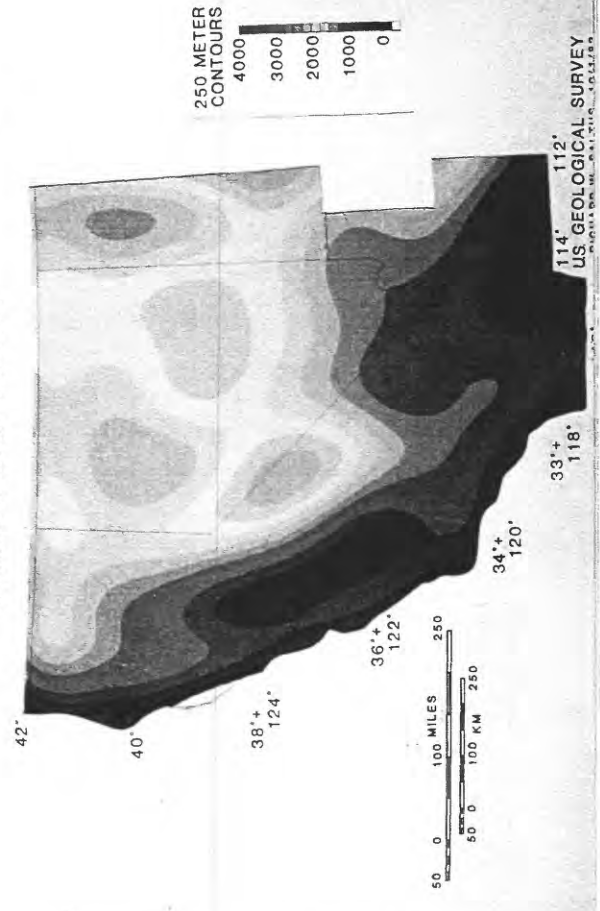


Figure 4

Figure 5

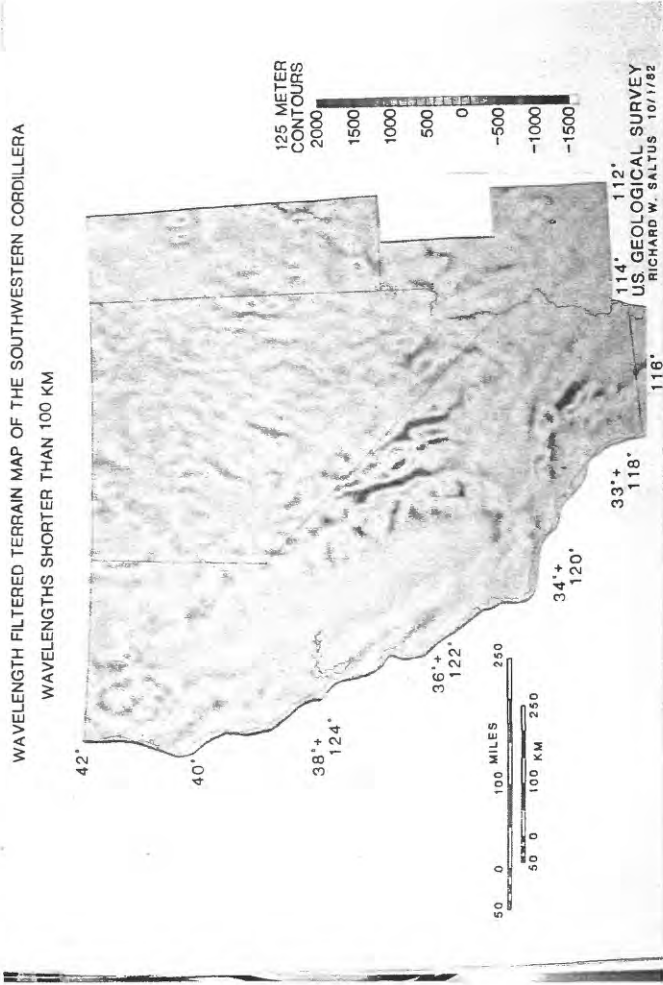


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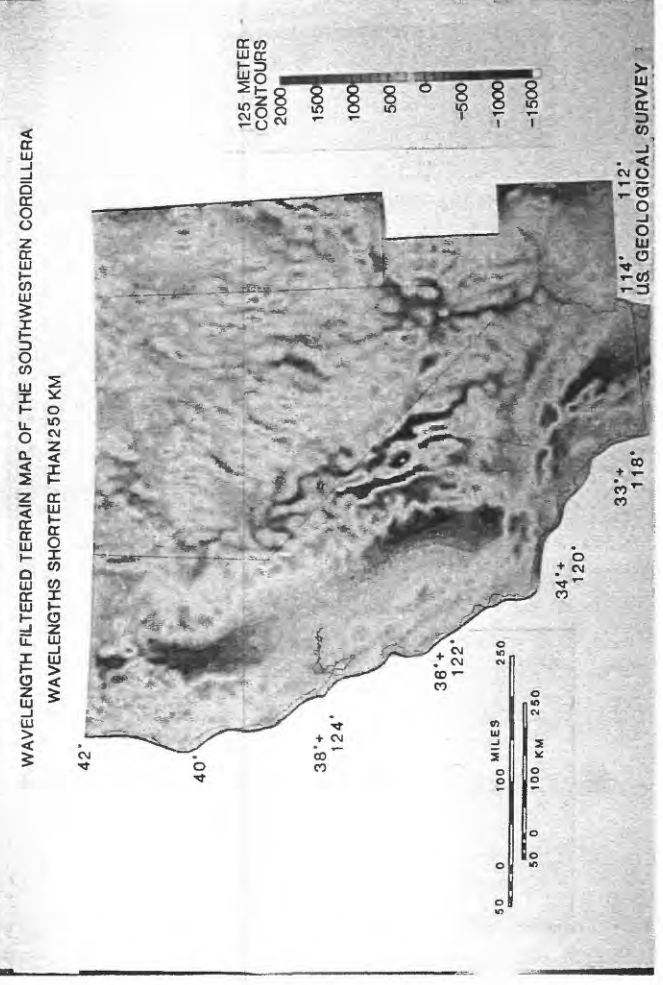
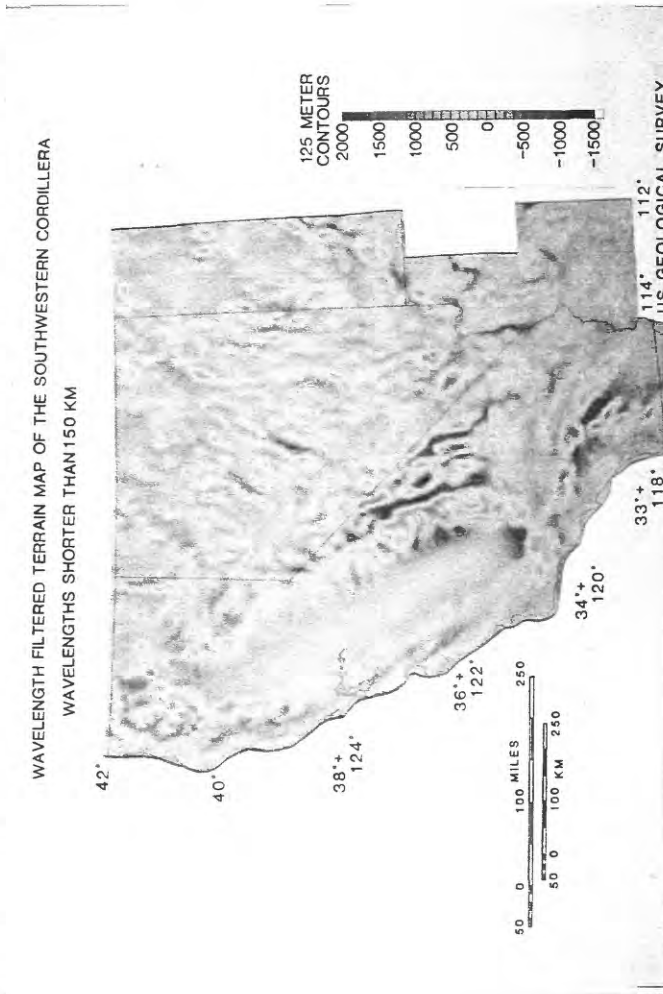


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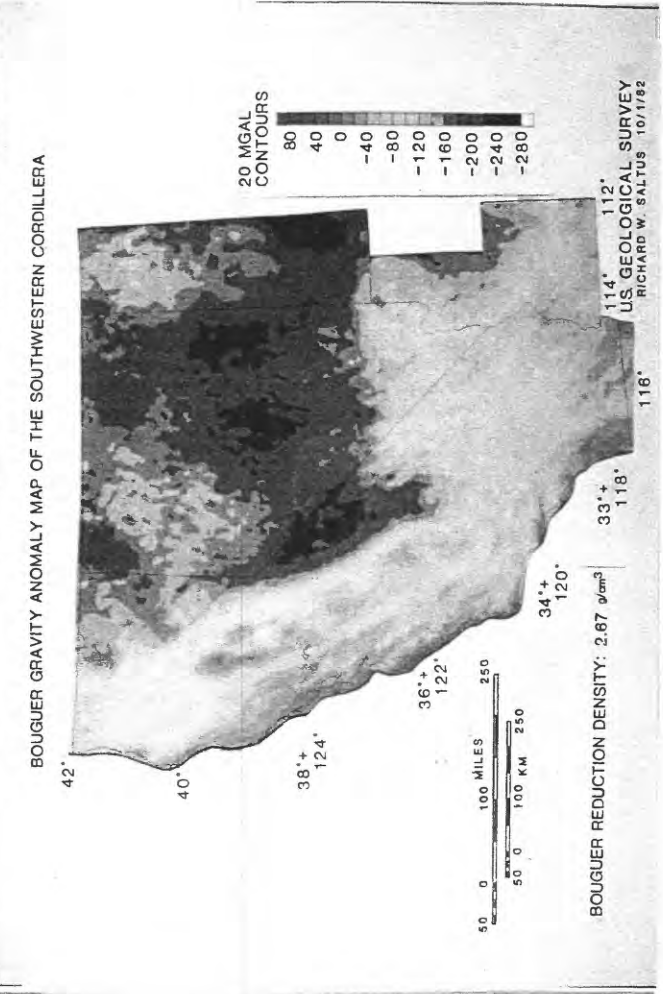


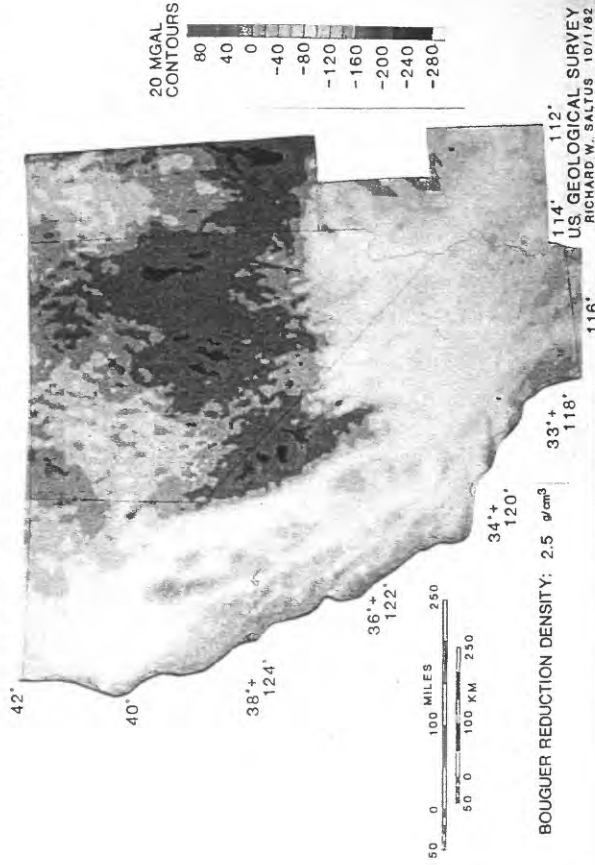
Figure 8

Figure 9

USGS Open-file 84-95

Figure 10

BOUGUER GRAVITY ANOMALY MAP OF THE SOUTHWESTERN CORDILLERA



ISOSTATIC RESIDUAL GRAVITY MAP OF THE SOUTHWESTERN CORDILLERA

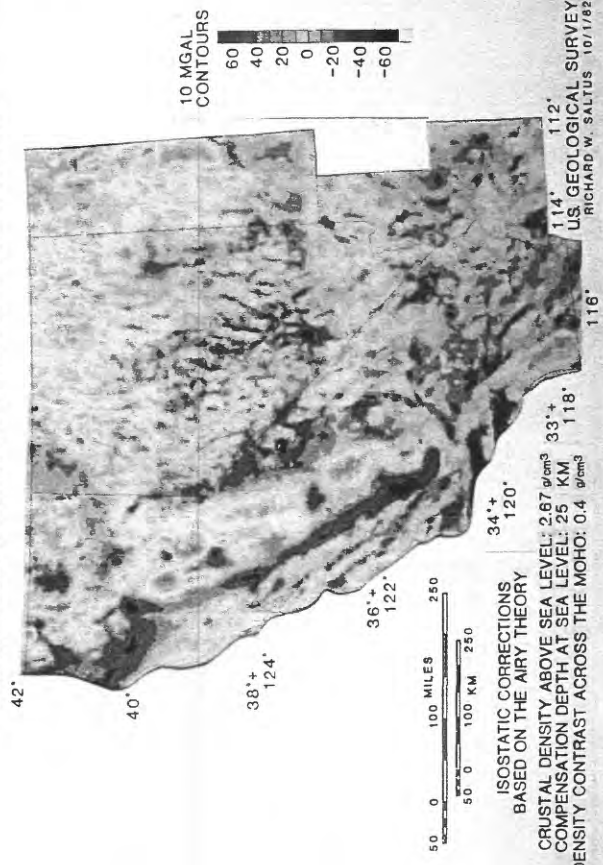
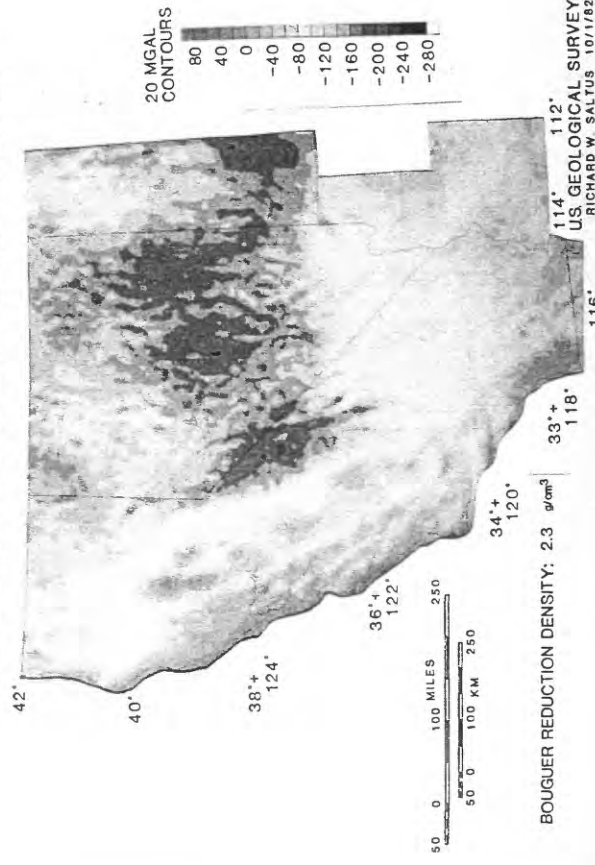


Figure 11

BOUGUER GRAVITY ANOMALY MAP OF THE SOUTHWESTERN CORDILLERA



ISOSTATIC RESIDUAL GRAVITY MAP OF THE SOUTHWESTERN CORDILLERA

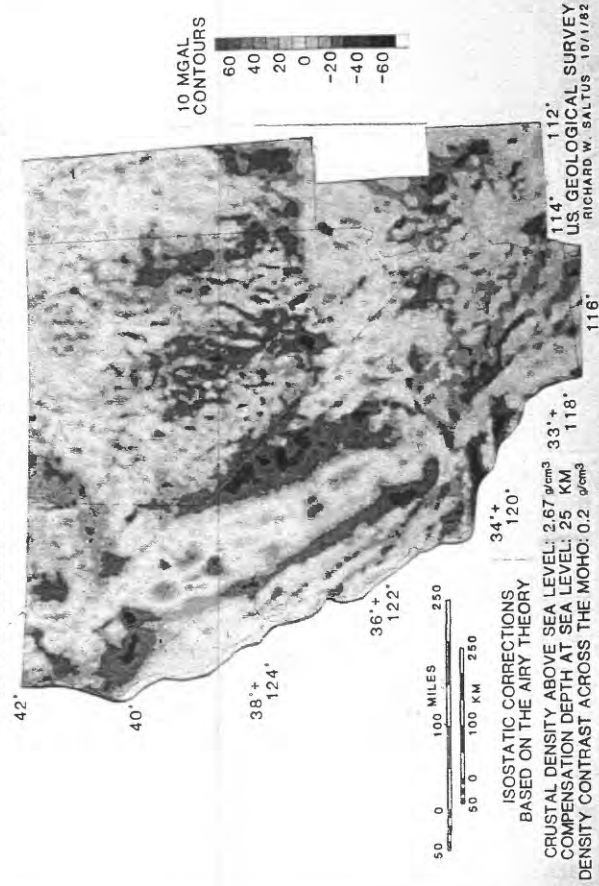


Figure 12

Figure 13

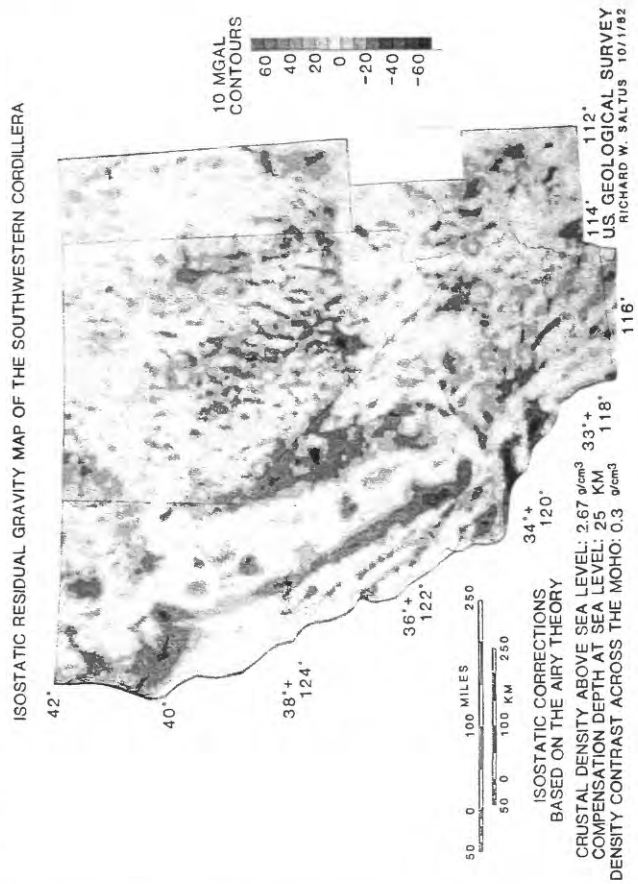


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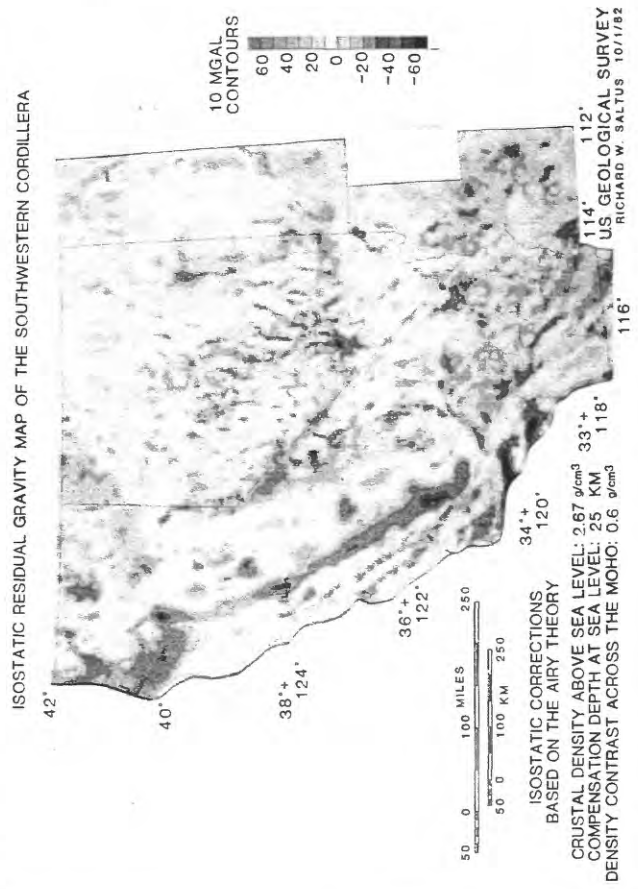


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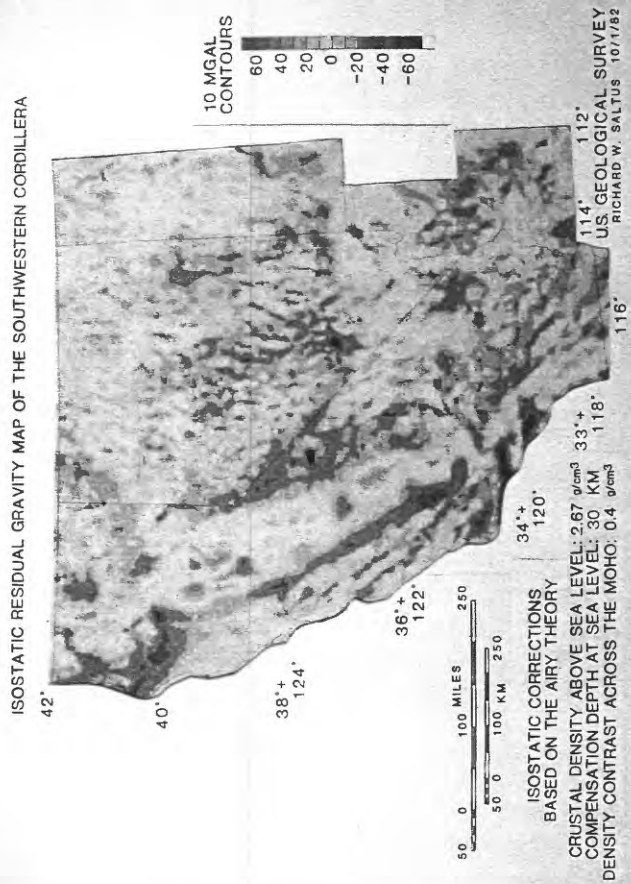


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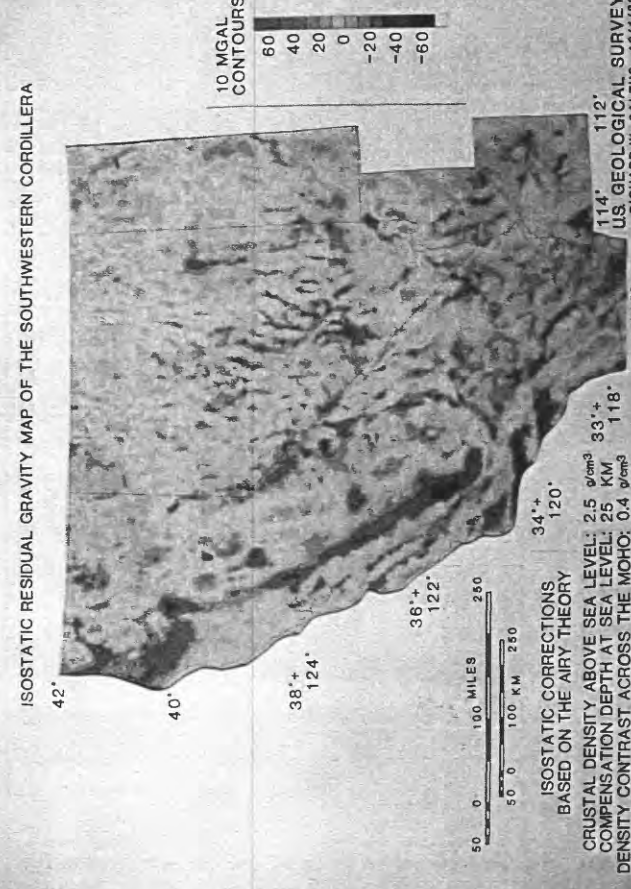


Figure 17

ISOSTATIC RESIDUAL GRAVITY MAP OF THE SOUTHWESTERN CORDILLERA

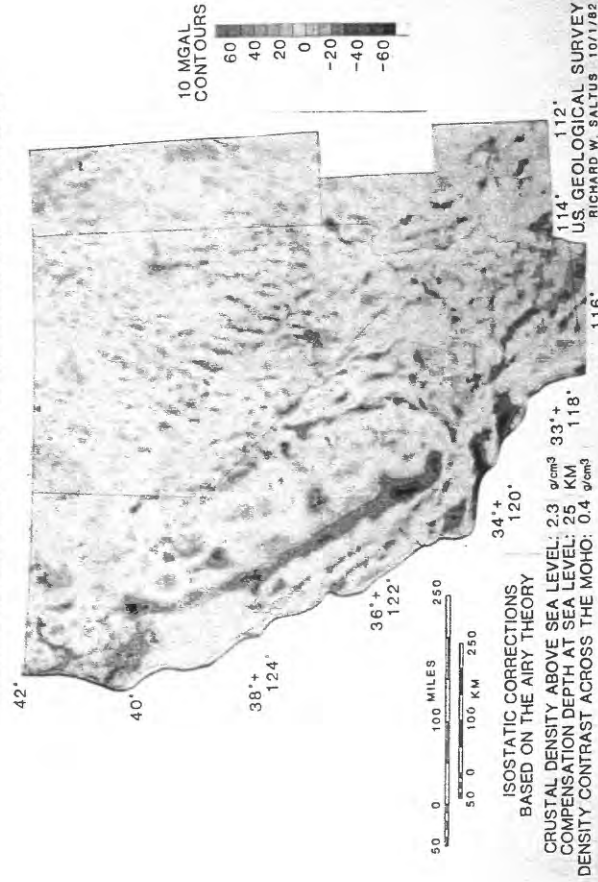
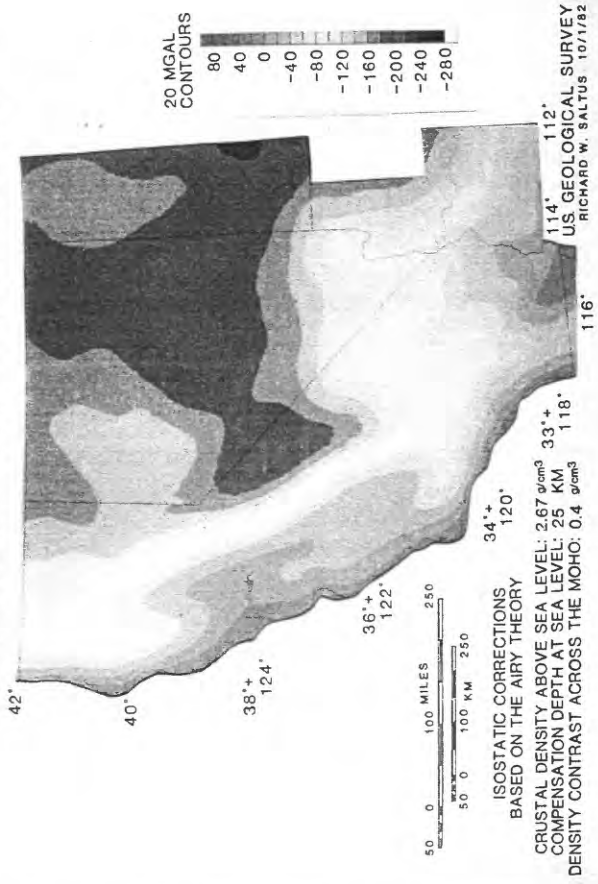


Figure 18

USGS Open-file 84-95

ISOSTATIC REGIONAL GRAVITY MAP OF THE SOUTHWESTERN CORDILLERA



ISOSTATIC REGIONAL GRAVITY MAP OF THE SOUTHWESTERN CORDILLERA

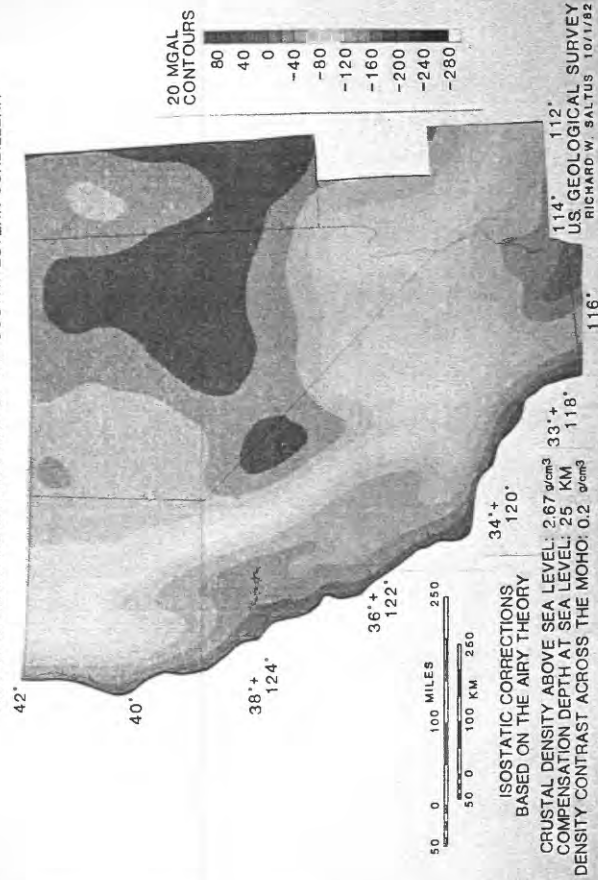


Figure 19

ISOSTATIC REGIONAL GRAVITY MAP OF THE SOUTHWESTERN CORDILLERA

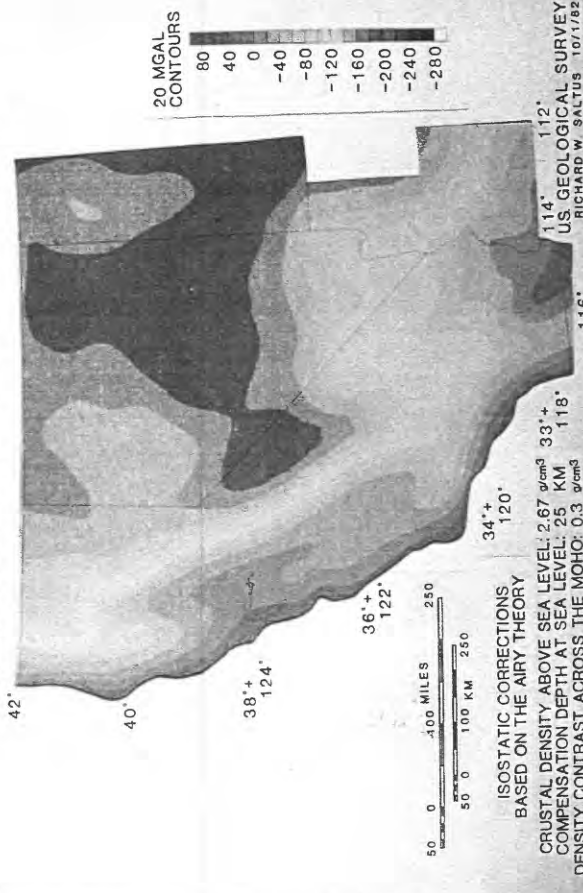


Figure 20

Figure 21

Figure 22

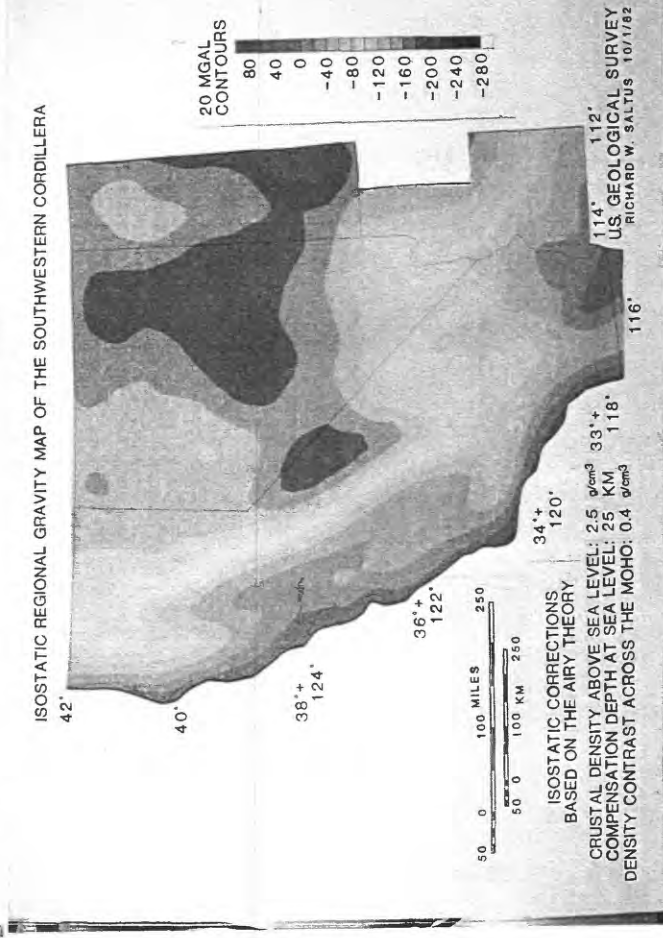
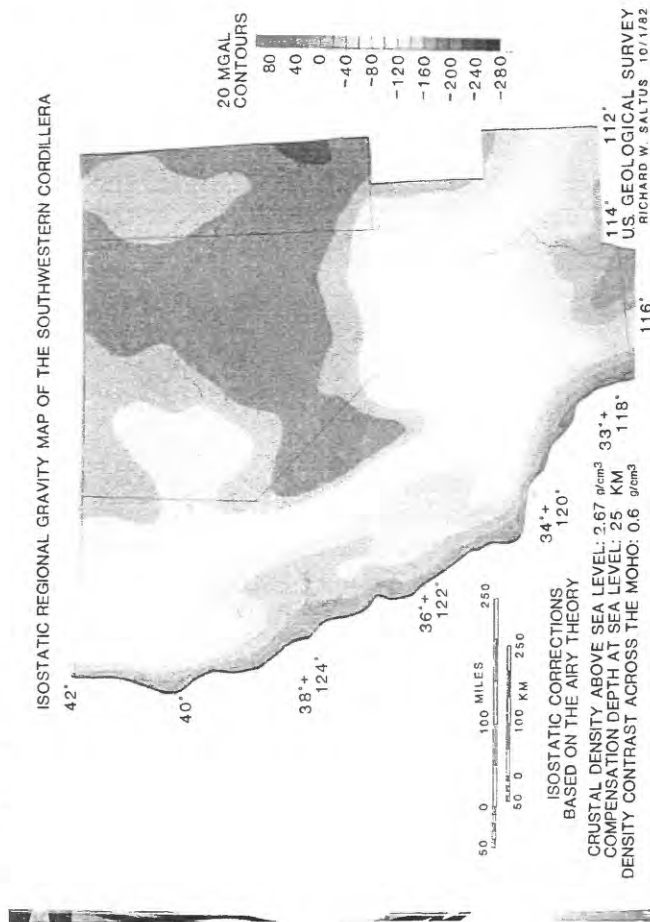


Figure 23

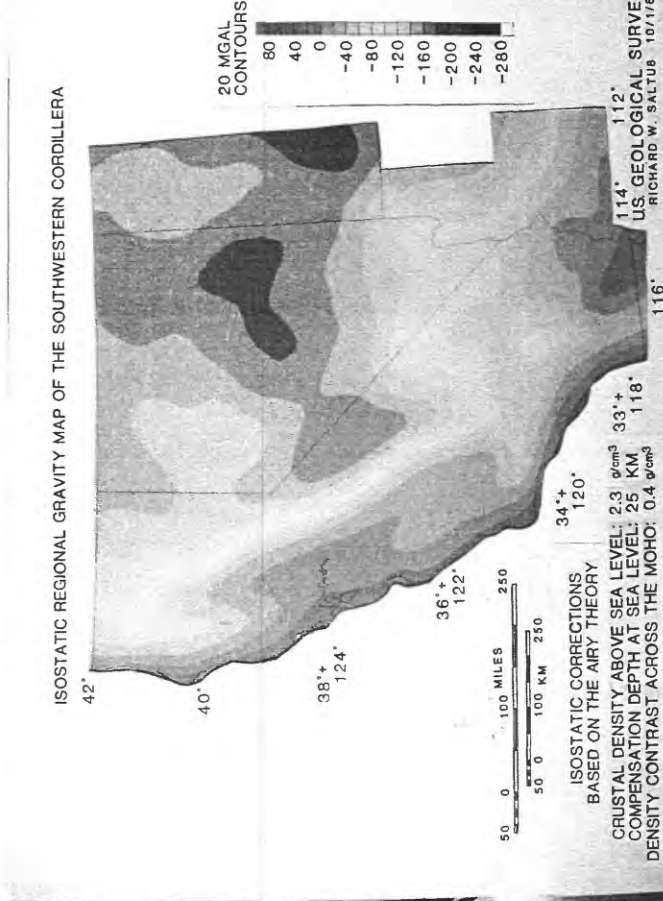
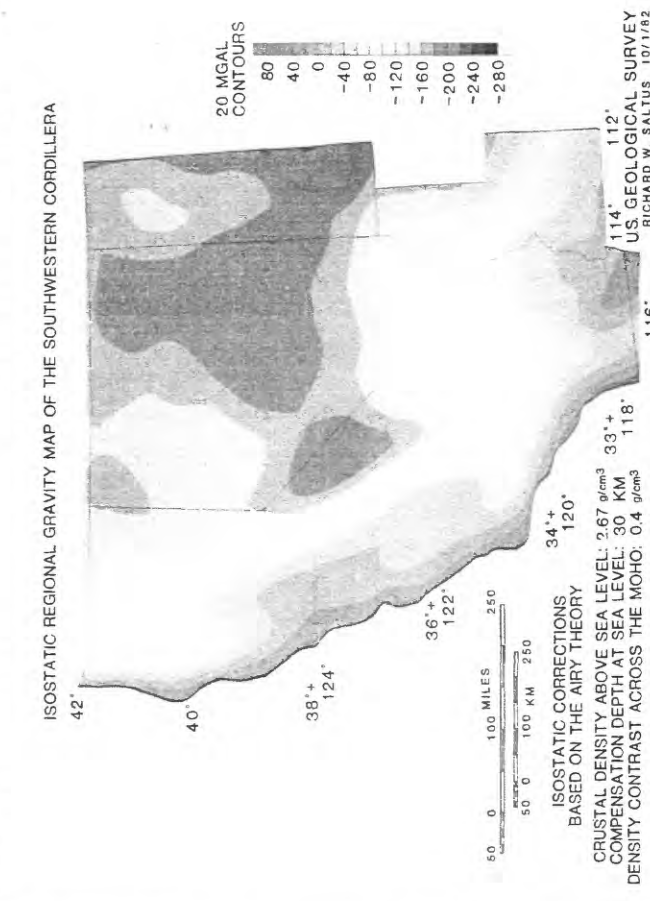


Figure 24

Figure 25

Figure 26

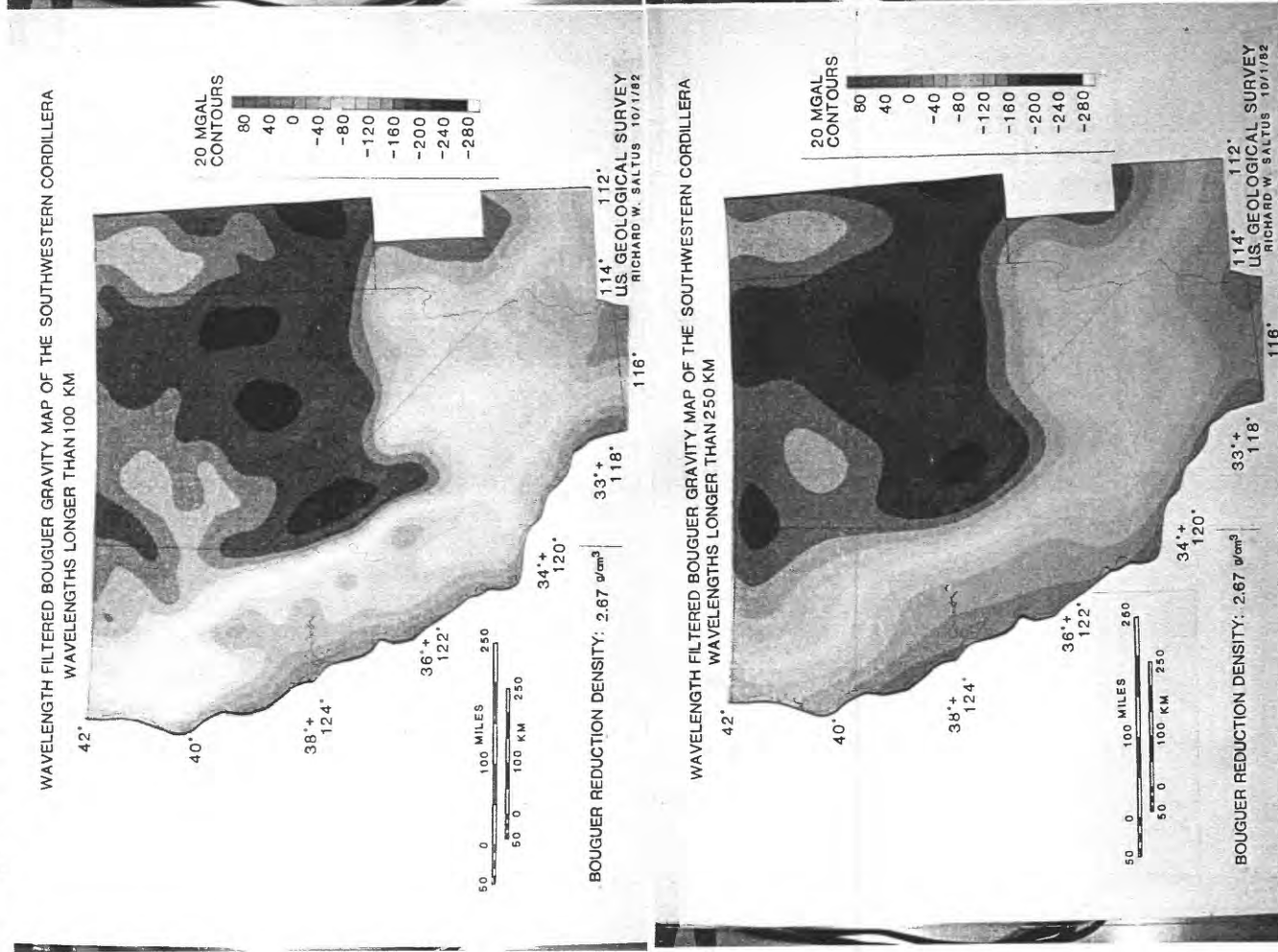


Figure 27

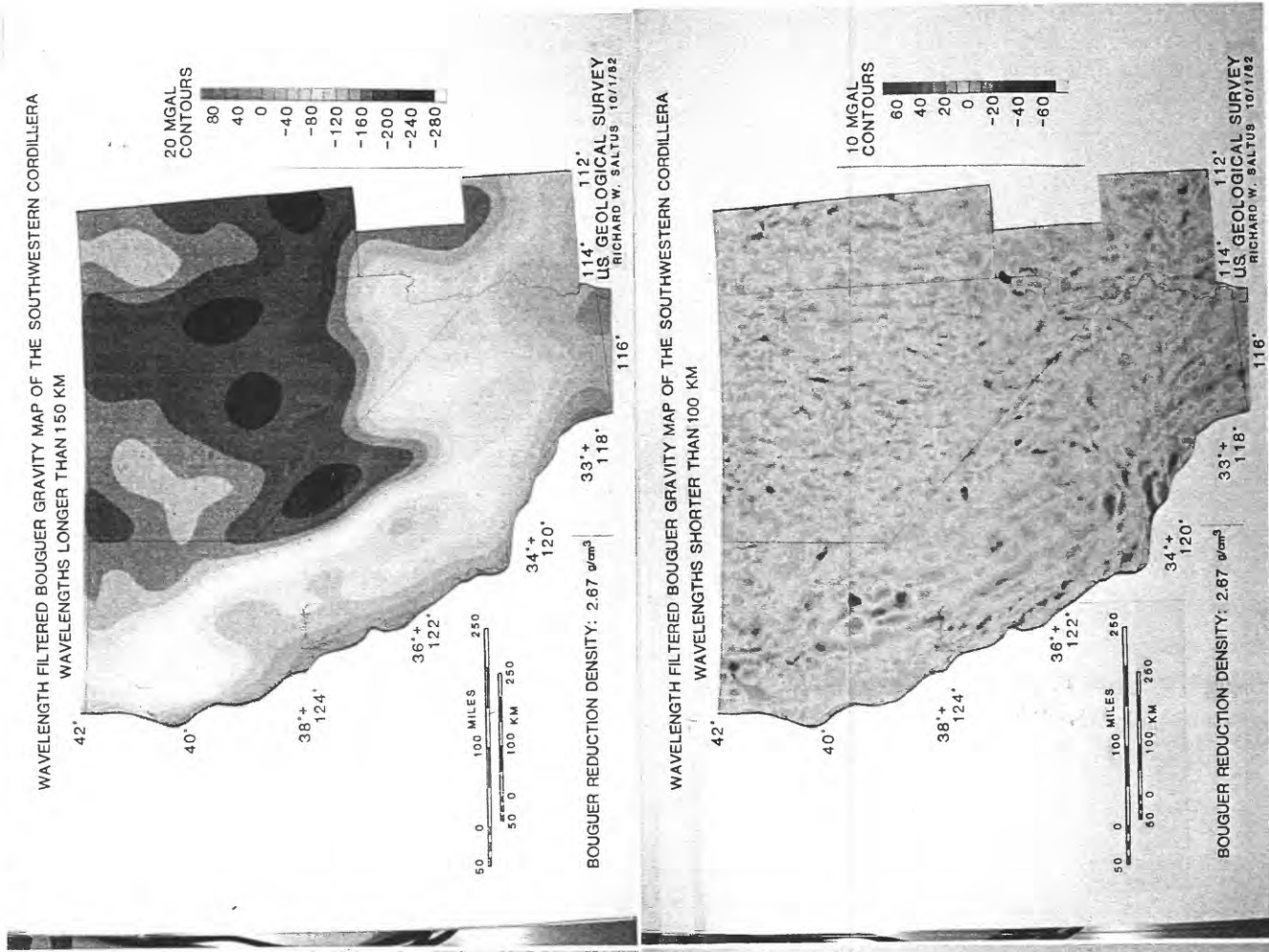


Figure 28

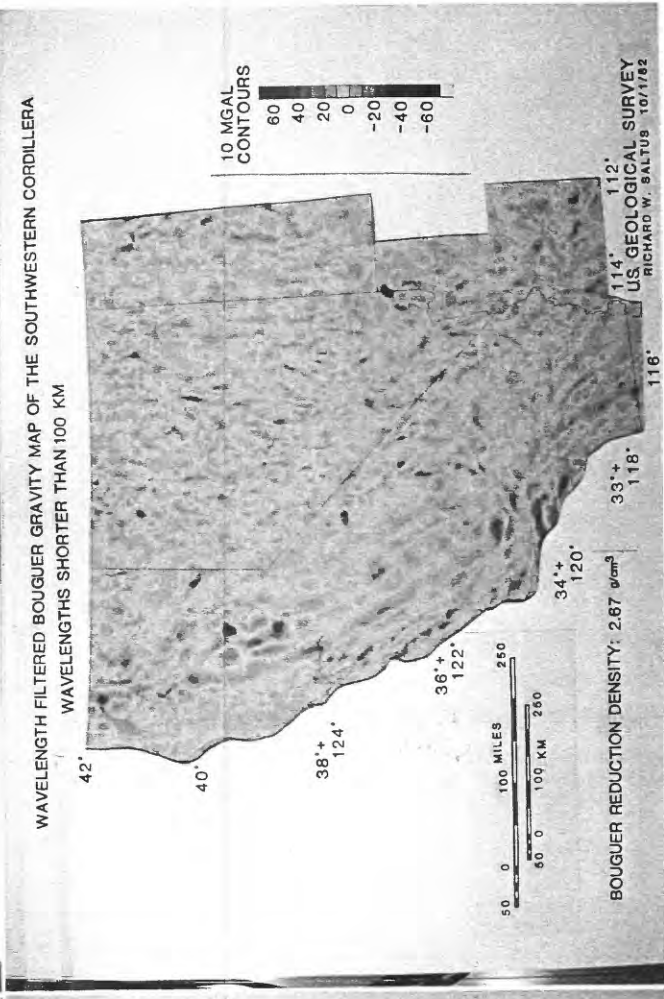


Figure 29

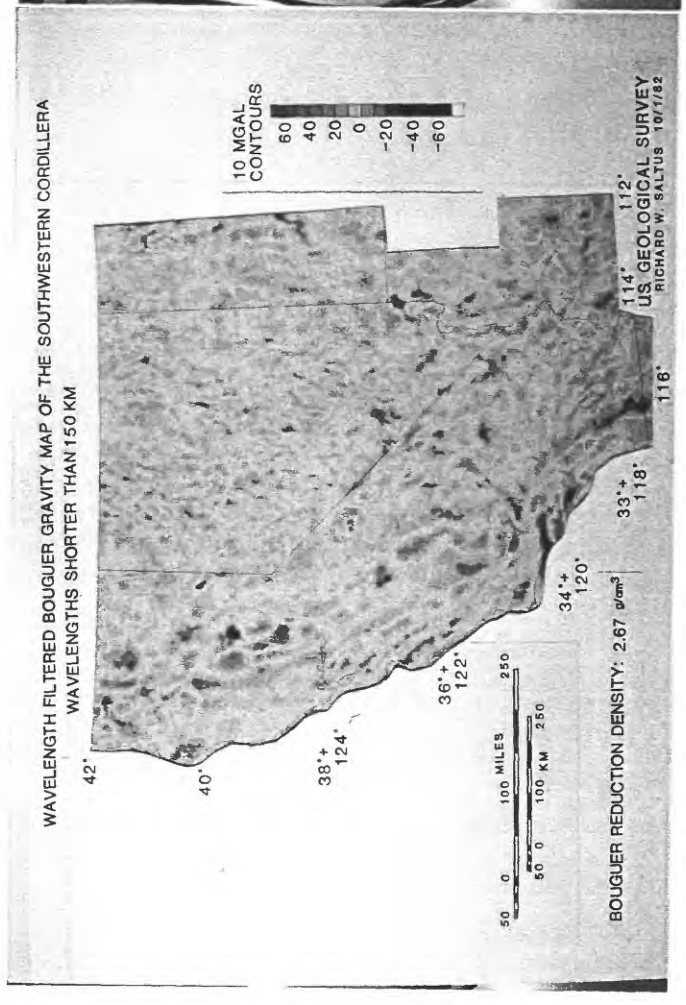


Figure 30

